

Towards GGOS Stations in New Zealand

Sergei Gulyaev

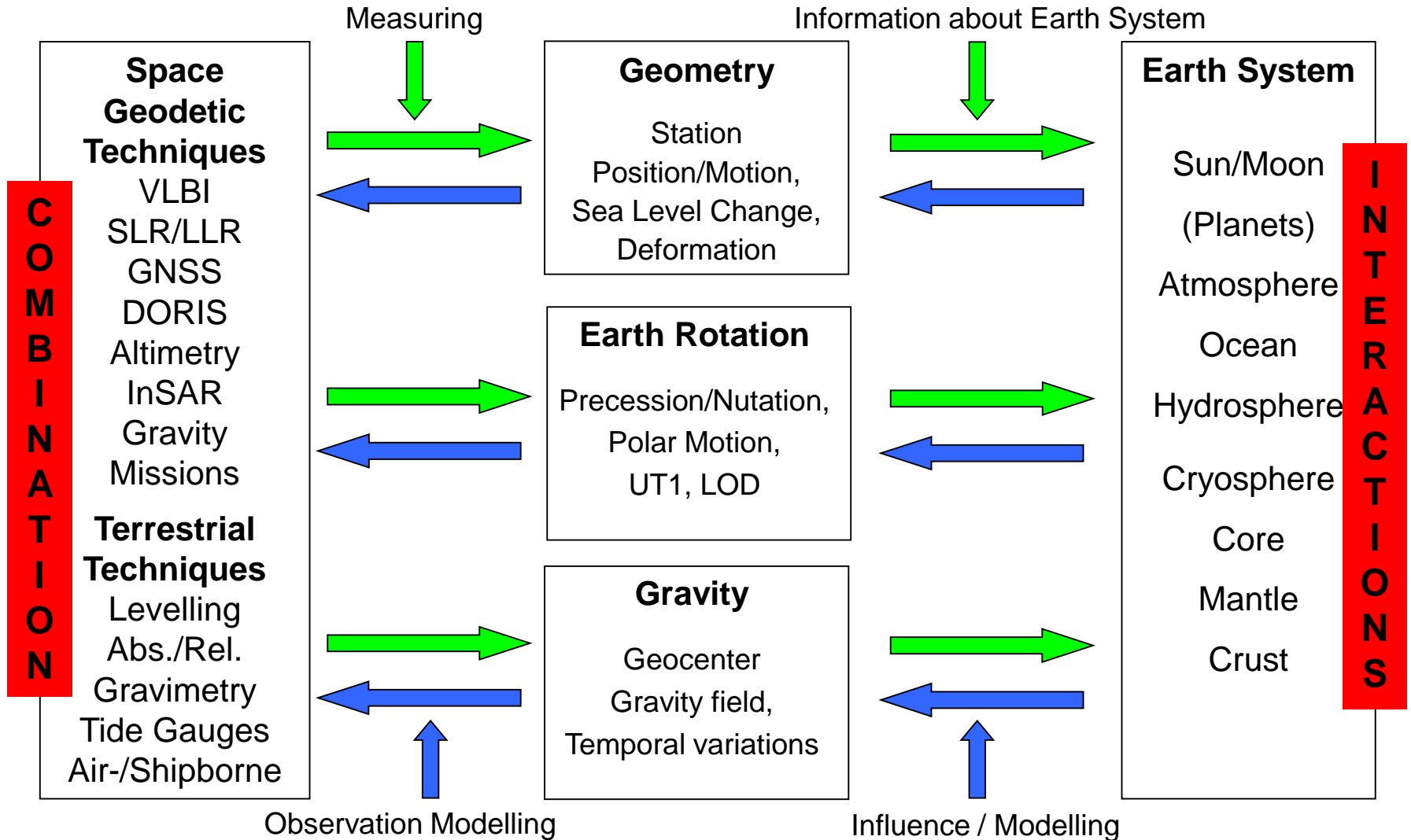
Institute for Radio Astronomy and Space Research
Auckland University of Technology

International Workshop “New Zealand Positional Strategy: GGOS Vision”
Land Information New Zealand, Wellington, 1 December 2014

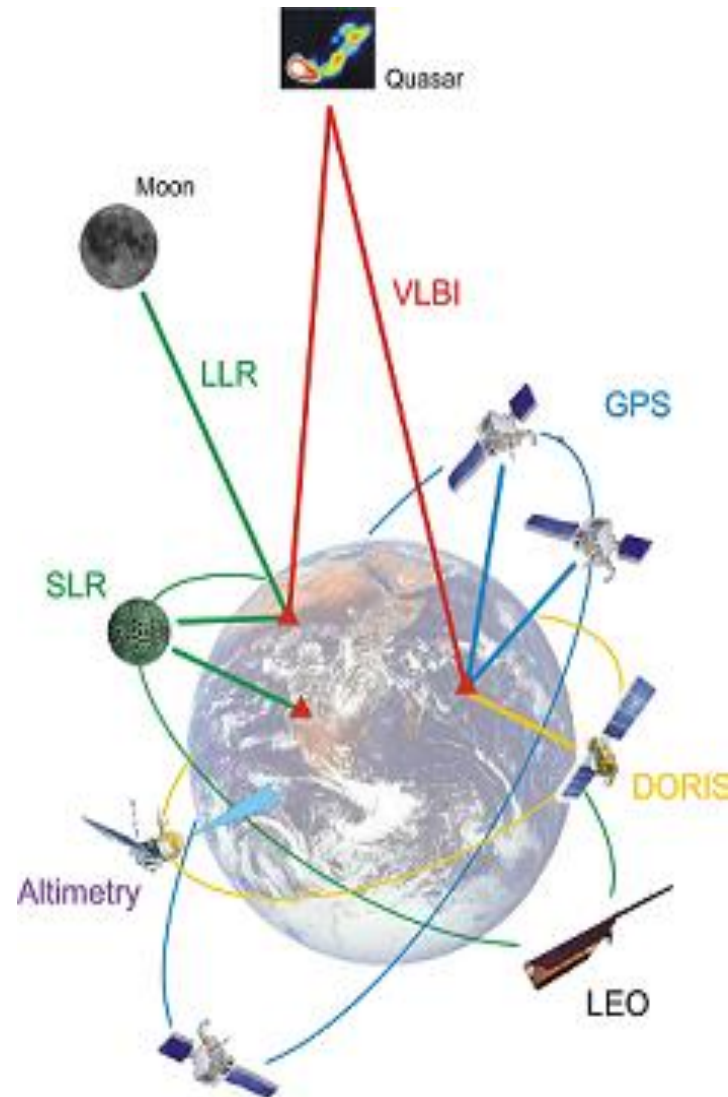
Space Geodesy is aimed

- To study Earth motions (rotation, irregularities of rotation, precession, nutation, polar motion, tectonic plate motion, intra-plate motions);
- To measure Earth parameters: scale (radius), position of the Earth centre (geocentre);
- To establish and support the International Celestial Reference Frame (VLBI only);
- To establish and support the International Terrestrial Reference Frame
- To determine coordinates of geodetic stations and variations of the coordinates and baselines (linear, seasonal signal, discontinuities).

Space Geodesy techniques



Space Geodesy techniques



Space Geodesy techniques

	Parameter Type	VLBI	GNSS	DORIS/ PRARE	SLR	LLR	Alti- metry	
ICRF	ICRF	X						Earth Rotation
	Nutation	X	(X)		(X)	X		
	Polar Motion	X	X	X	X	X		
	UT1	X						
	Length of Day (LOD)		X	X	X	X		
ITRF	Coord+Veloc (ITRF)	X	X	X	X	X	(X)	Gravity Field
	Geocenter		X	X	X		X	
	Gravity Field		X	X	X	(X)	X	
	Orbits		X	X	X	X	X	
Atmosphere	LEO Orbits		X	X	X		X	Atmosphere
	Ionosphere	X	X	X			X	
	Troposphere	X	X	X			X	
	Time/Freq.; Clocks	(X)	X		(X)			

(from Zuheir Altamimi, Presentation at the Fourth Meeting of the International Committee on GNSS, St. Petersburg, 14-18 Sept, 2009)

Space Geodesy techniques

Mix of techniques is fundamental to realize a frame (ITRF) that is stable in origin, scale, and with sufficient coverage

Technique Signal Source Obs. Type	VLBI Microwave Quasars Time difference	SLR Optical Satellite Two-way absolute range	GPS Microwave Satellites Range change	DORIS
Celestial Frame & UT1	Yes	No	No	No
Polar Motion	Yes	Yes	Yes	Yes
Scale	Yes	Yes	No (but maybe in the future!)	Yes
Geocenter ITRF Origin	No	Yes	Future	Future
Geographic Density	No	No	Yes	Yes
Real-time & ITRF access	Yes	Yes	Yes	Yes
Decadal Stability	Yes	Yes	Yes	Yes

(from Zuheir Altamimi, Presentation at the Fourth Meeting of the International Committee on GNSS, St. Petersburg, 14-18 Sept, 2009)

GGOS

- GGOS aims:
 - by using and integrating advanced geodetic observing techniques
 - 1) accuracy of < **1 mm** for position
 - 2) accuracy of < **0.1 mm/yr** for velocity

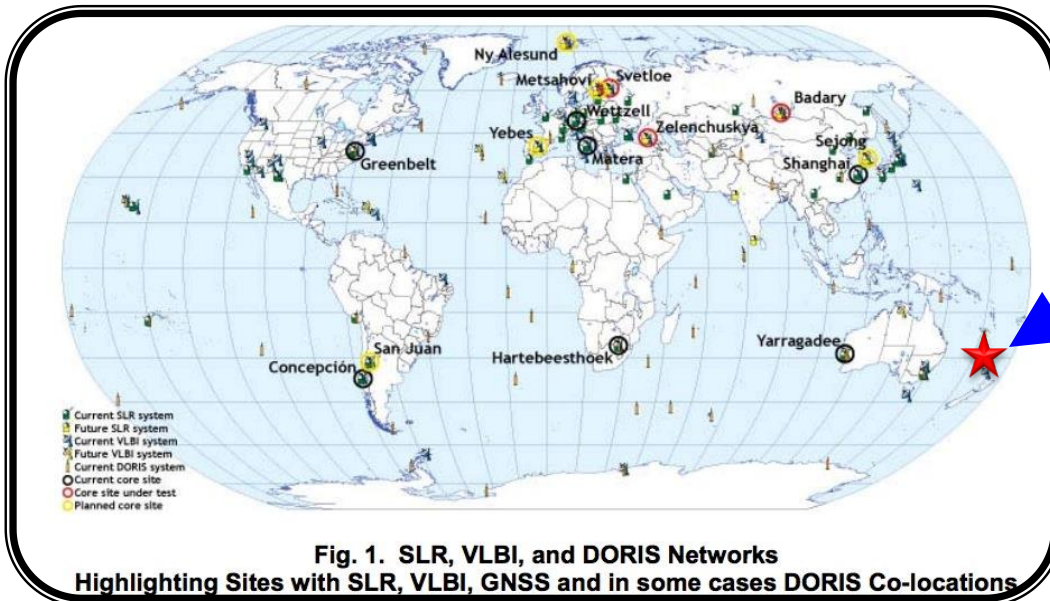
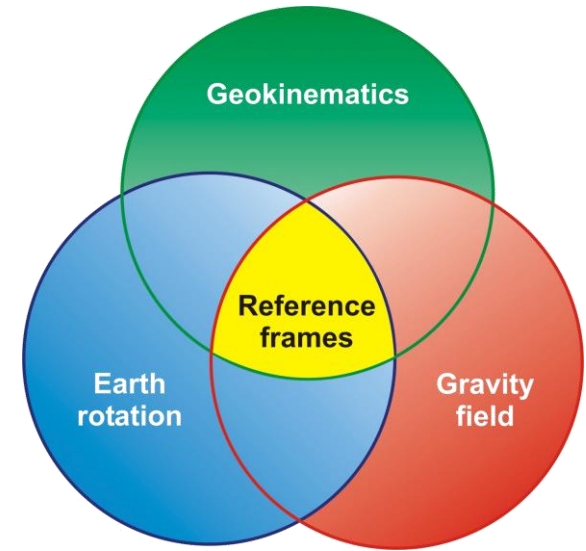


Fig. 1. SLR, VLBI, and DORIS Networks

Highlighting Sites with SLR, VLBI, GNSS and in some cases DORIS Co-locations

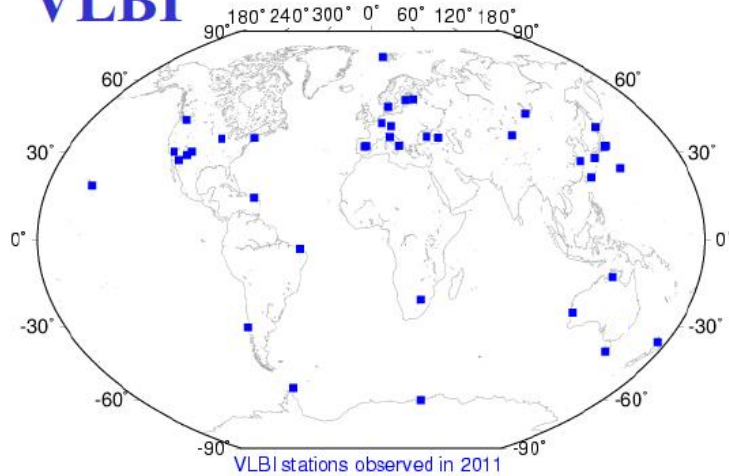
Warkworth



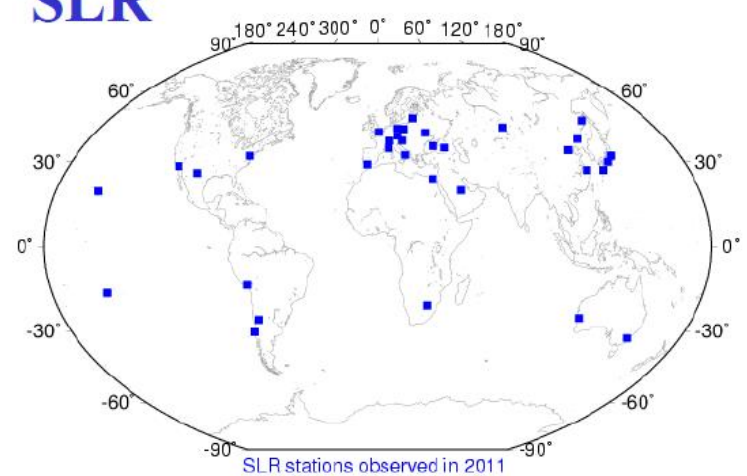
has capability to become a core site
Important contribution to geometrical distribution

GGOS

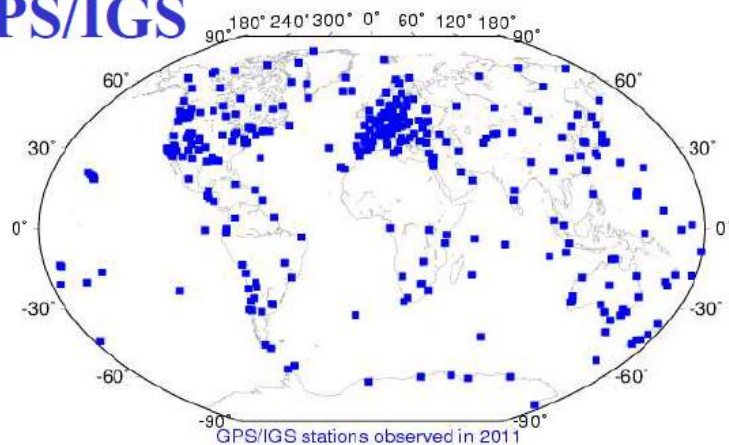
VLBI



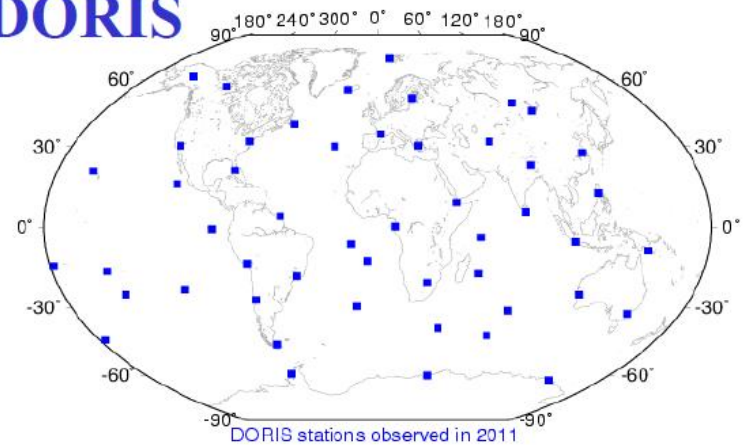
SLR



GPS/IGS



DORIS

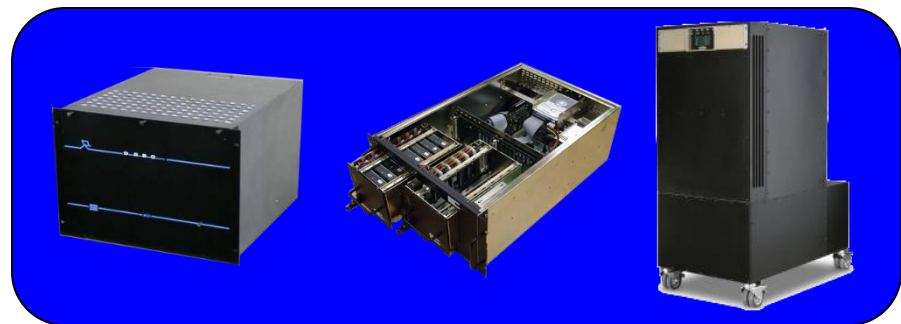


Geodetic 12m RT

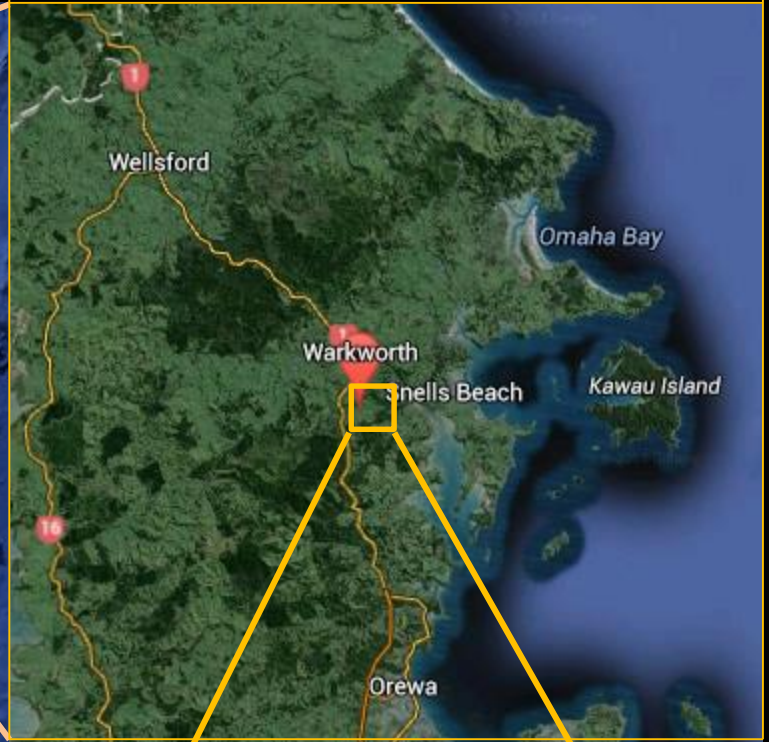
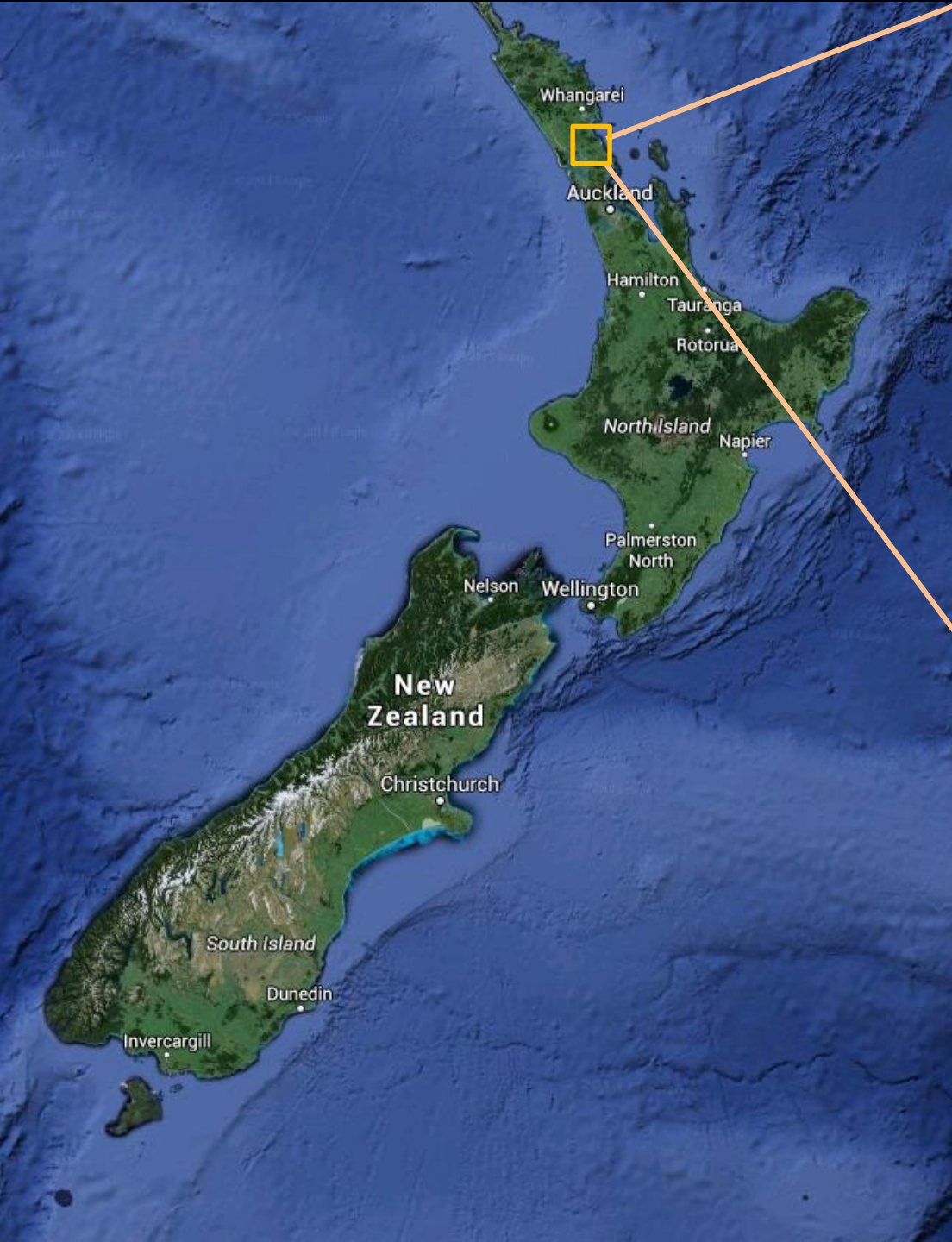
- Warkworth 12m radio telescope
 - Launched on October, 2008
 - Participates in the IVS sessions since the beginning of 2011



WARK12M









Satellite Station Rd



The 30m RT at
Warkworth

Thanks to Telecom

Cost of new 30m RT
would be US\$25-30m













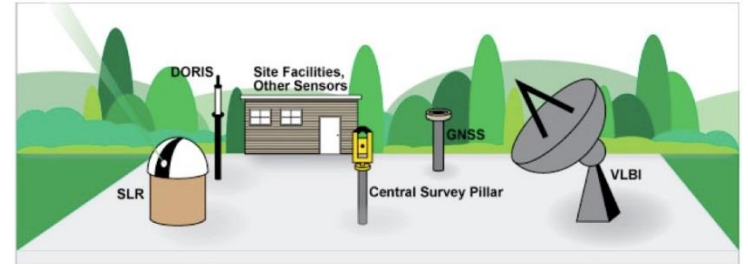








To make a GGOS core site



✧ Stable results

✧ Make an appropriate model

- ✧ atmospheric delay
- ✧ ocean tide loading
- ✧ environmental loads, etc

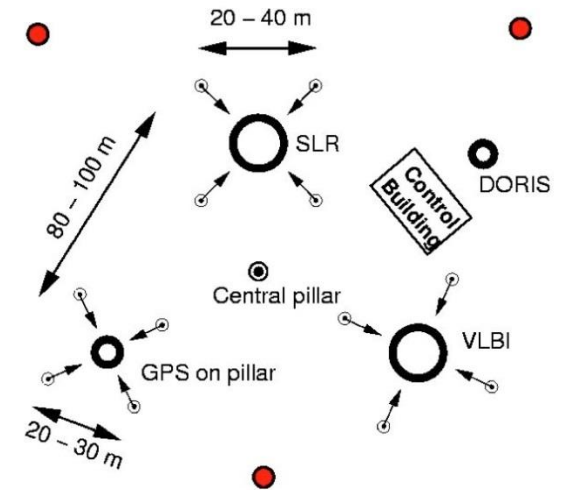
✧ assistance observation

- ✧ Groundwater,
Soil moisture, etc.

✧ Cooperation with other institutes

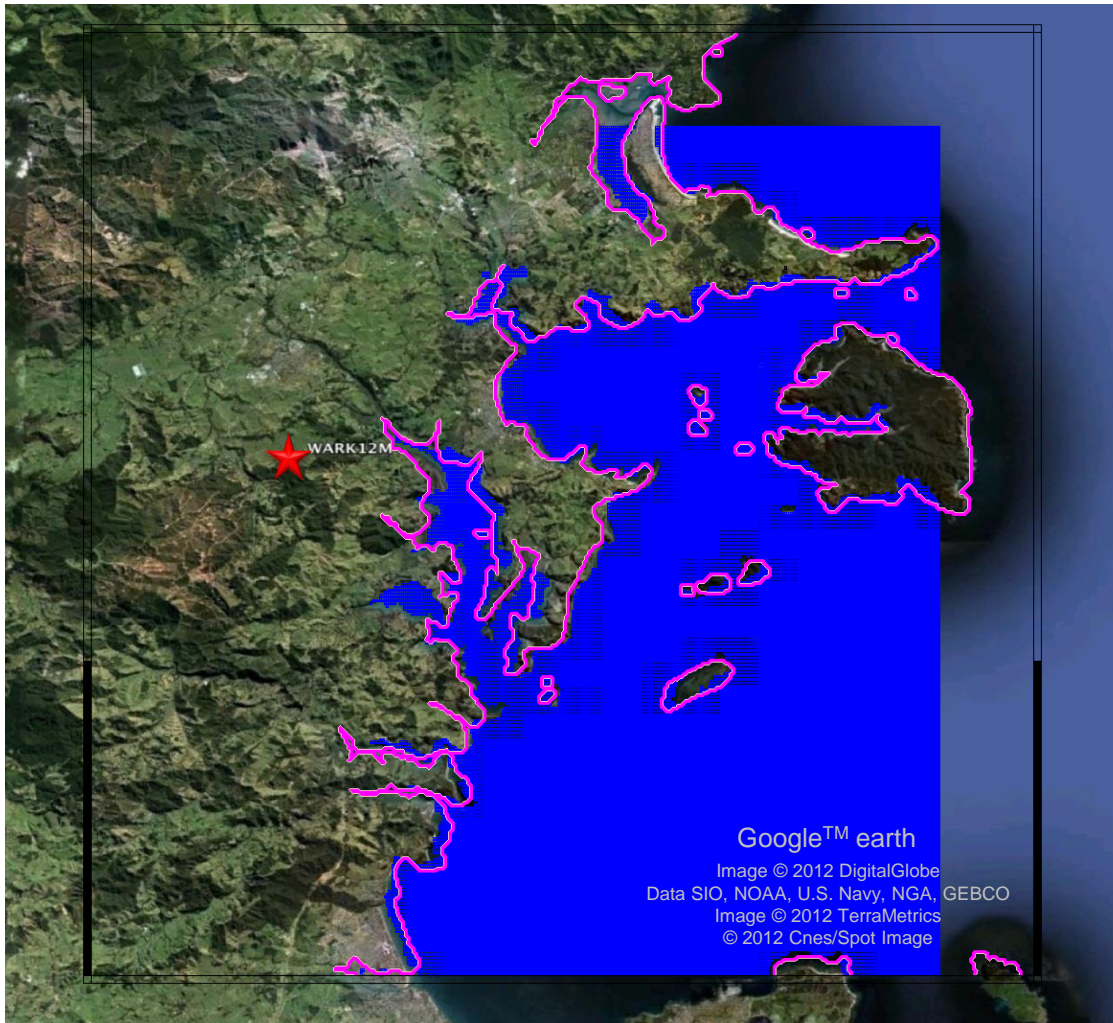
✧ periodical local survey

- ✧ to realize the local tie between VLBI and GPS
- ✧ LINZ and GNS Science



○ Survey pillars with unobstructed views
“Site Requirements for GGOS Core Sites”

Land-sea data for Warkworth



full resolution coastline
of GMT package

Shuttle Radar
Topography Mission
The Mission to Map the World

Version 2 : SRTM3
World 3 arc-seconds
(90m)



The 2012 Warkworth Observatory Local Tie Survey

The 2012 Warkworth Observatory Local Tie Survey



Land Information New Zealand
Record A1387321



Azimuth Observations (5 targets)

The elevation axis was fixed in a vertical setting. The azimuth axis was rotated in 20 degree increments starting from 340 degrees backward through to 0 degrees. One set of observations was completed to all visible targets on the radio telescope for each 20 degree rotation.



Figure 6.2: Target placement on the front (left and top right) and back (bottom right) on telescope for the azimuth observations.

Elevation Observation (4 targets)

The azimuth axis was fixed in a horizontal orientation of 13 degrees and 130 degrees, to be orthogonal to the total station line of sight. The elevation axis was set to 85 (due to the bearing issue could not be set to 90) and lowered in 10 degree increments down to 10 degrees.



Figure 6.3: Placement of targets on dish for the elevation observations

Gravity

AG

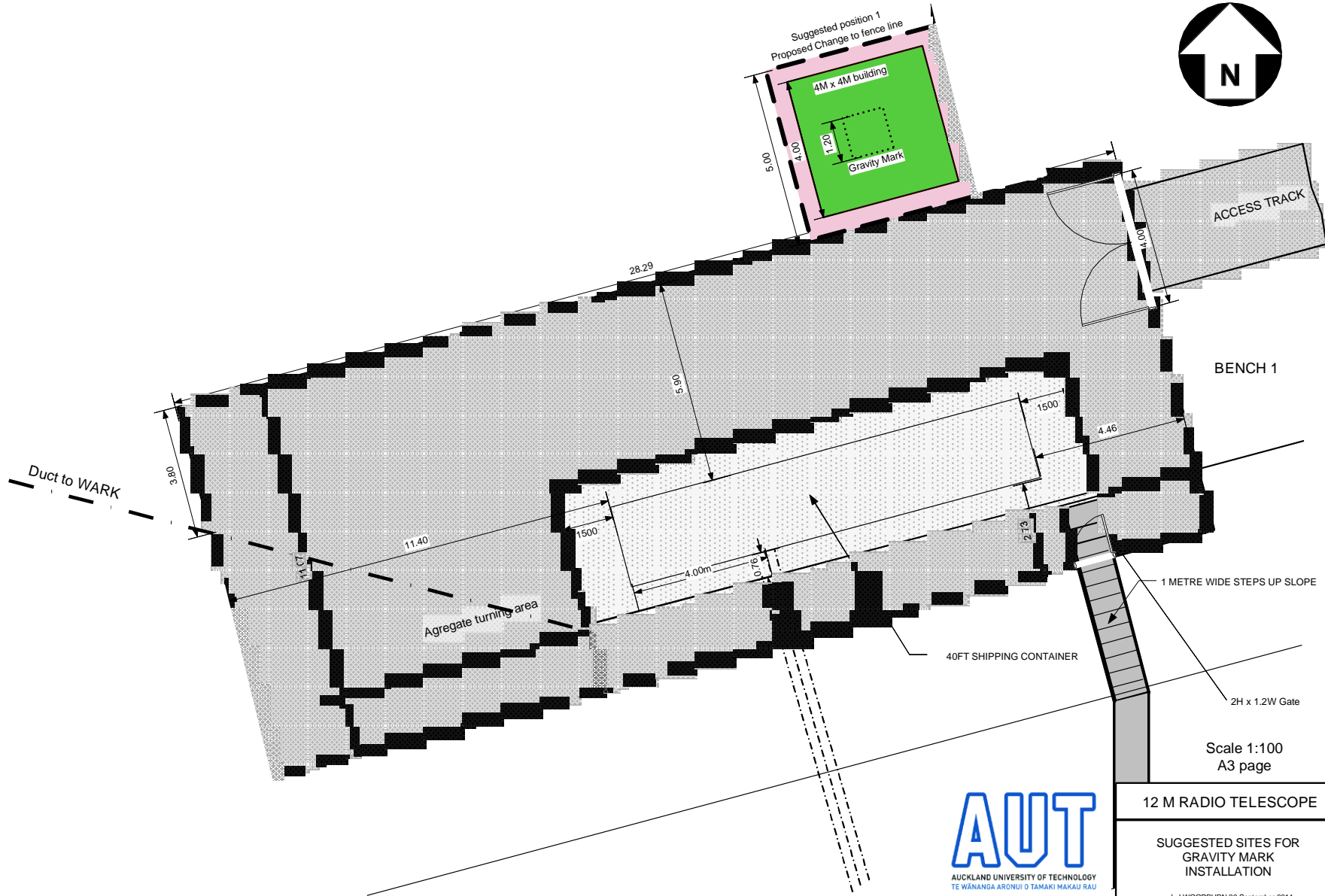


<http://www.micrograv.com/>

SG



<http://www.gwrinstruments.com/>



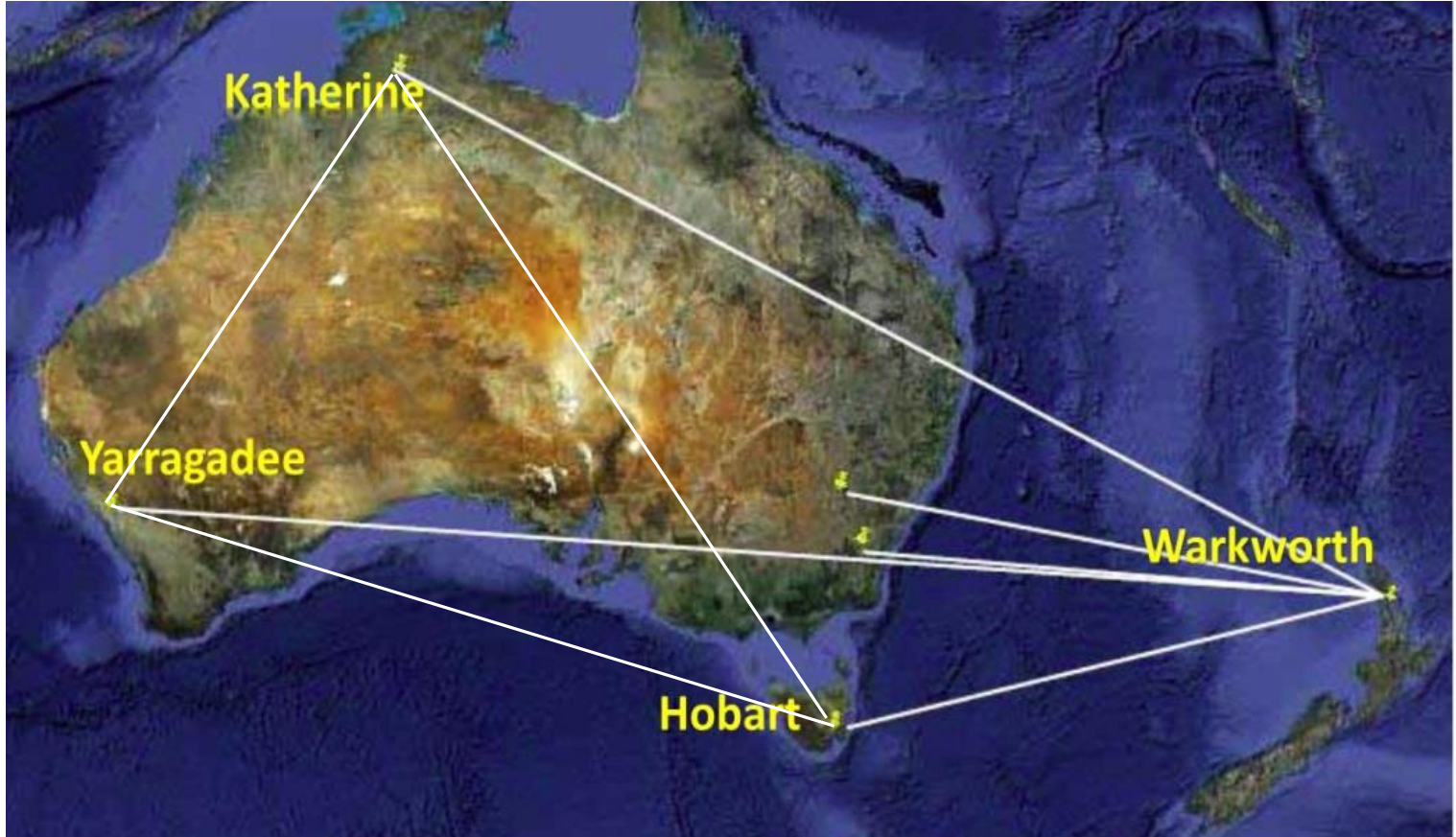
12 M RADIO TELESCOPE

SUGGESTED SITES FOR GRAVITY MARK INSTALLATION

AuScope



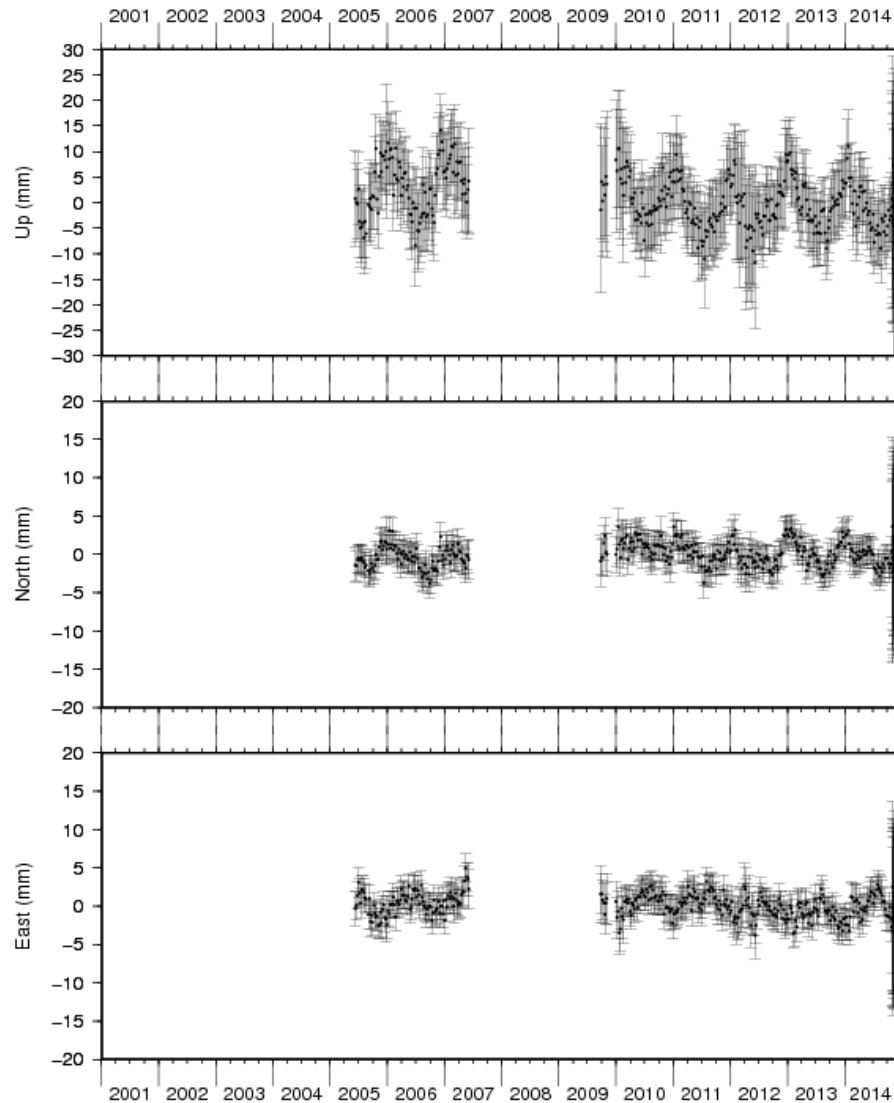
AuScope



GGOS stations in Australasia.

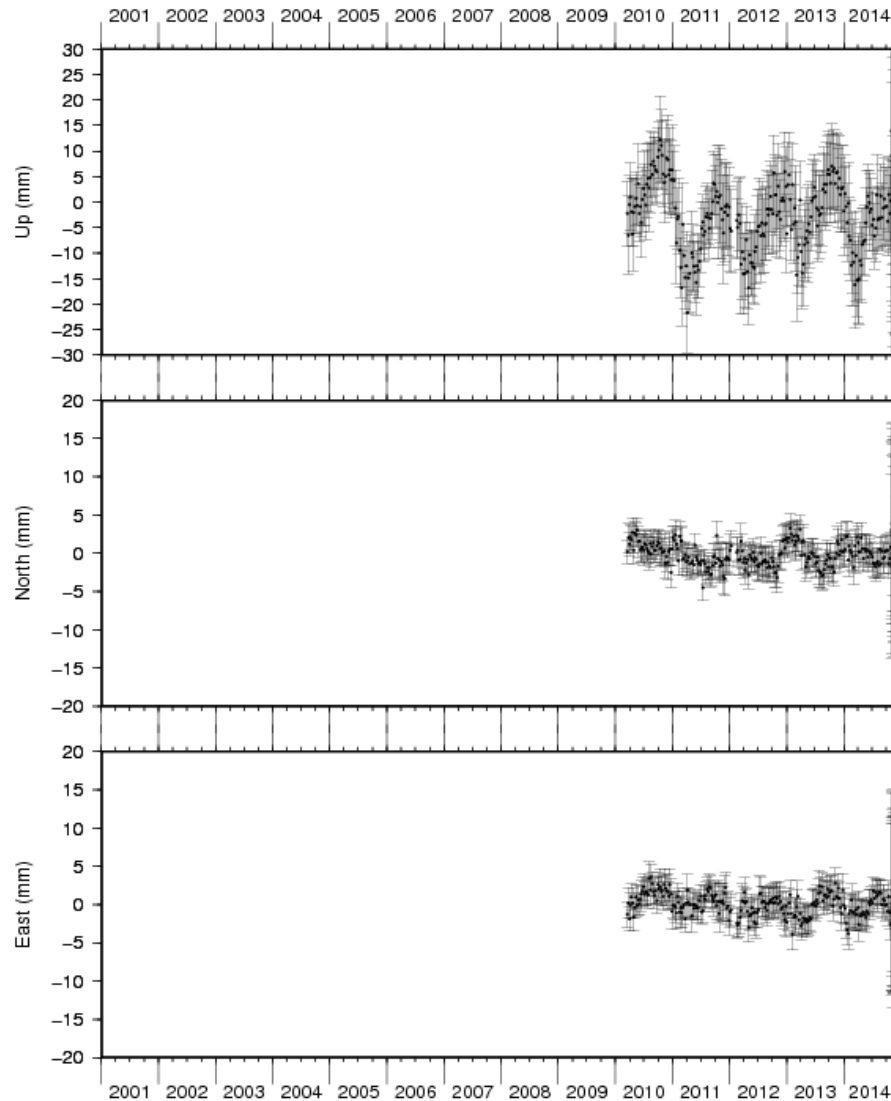
GNSS coordinate variations – YARR

YARR model residuals



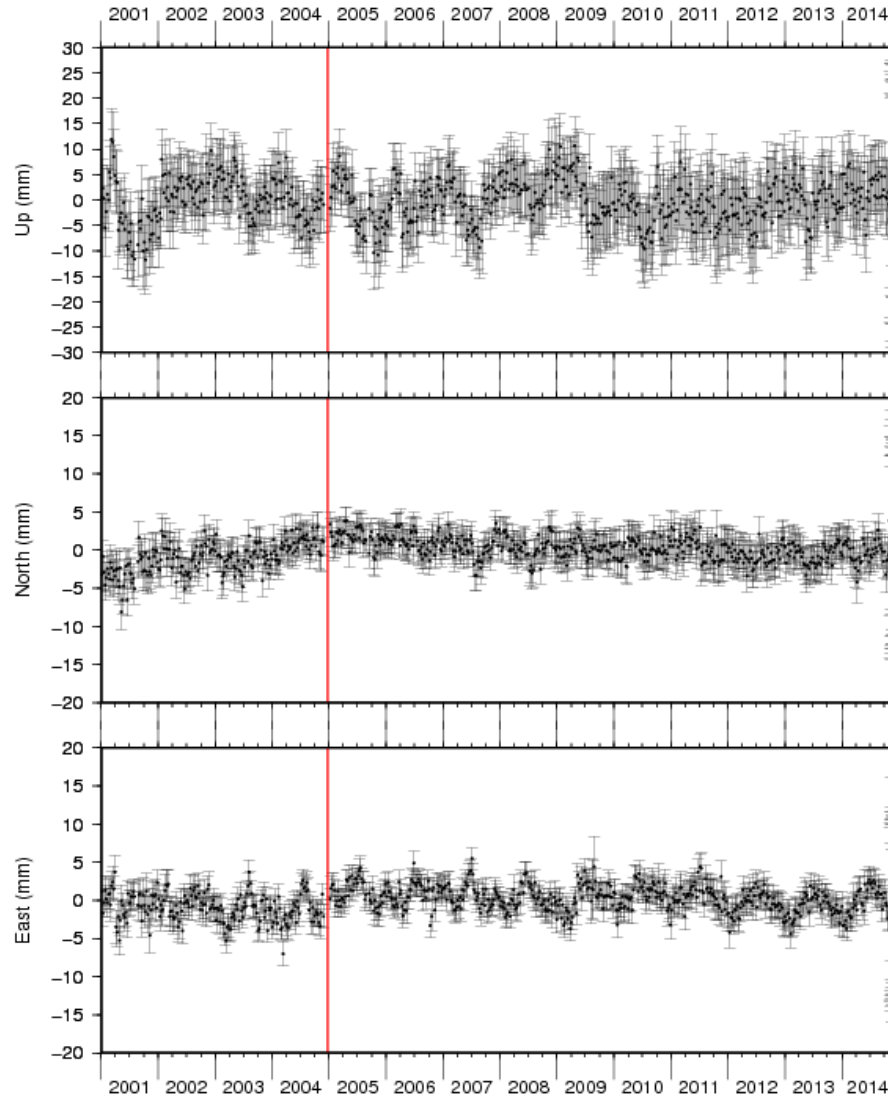
GNSS coordinate variations – KAT1

KAT1 model residuals



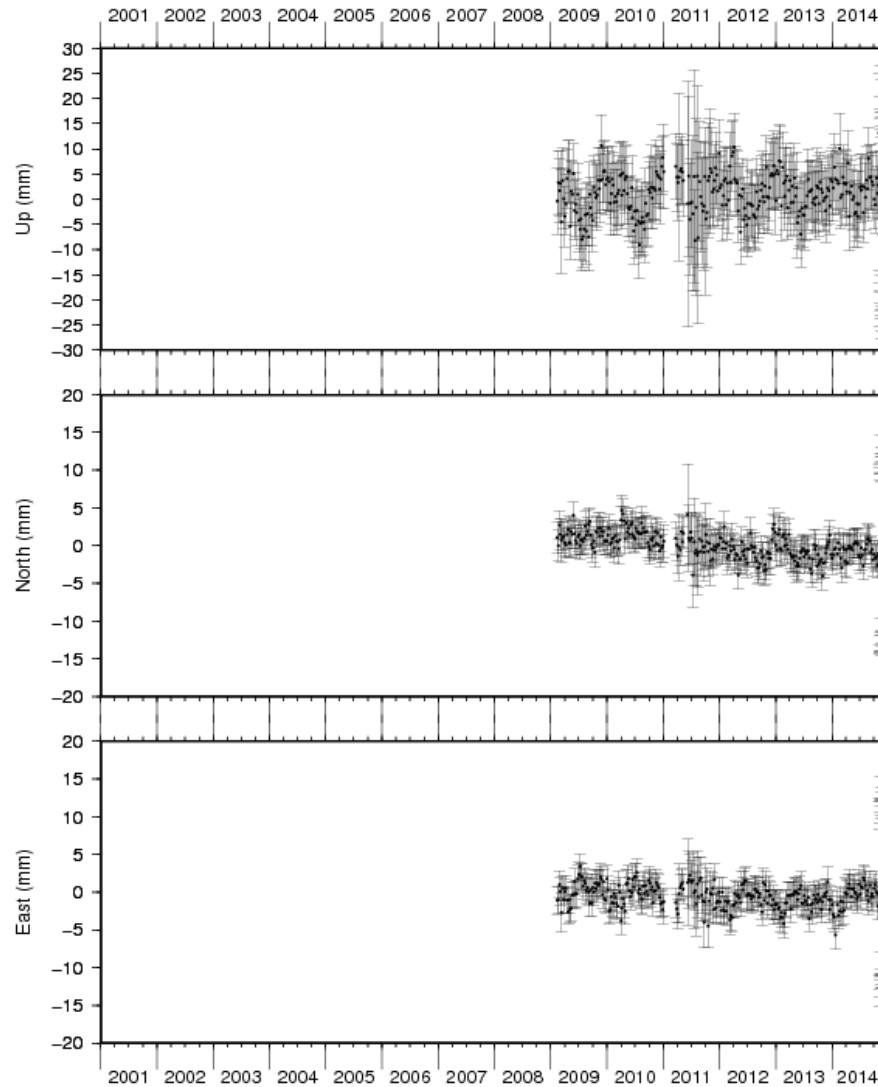
GNSS coordinate variations – HOB2

HOB2 model residuals

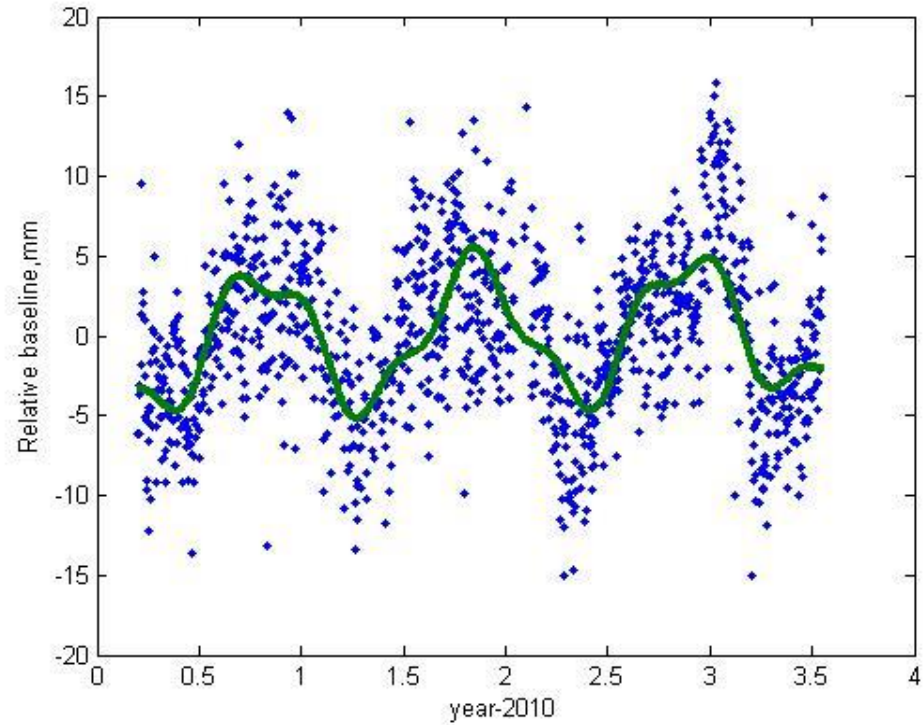
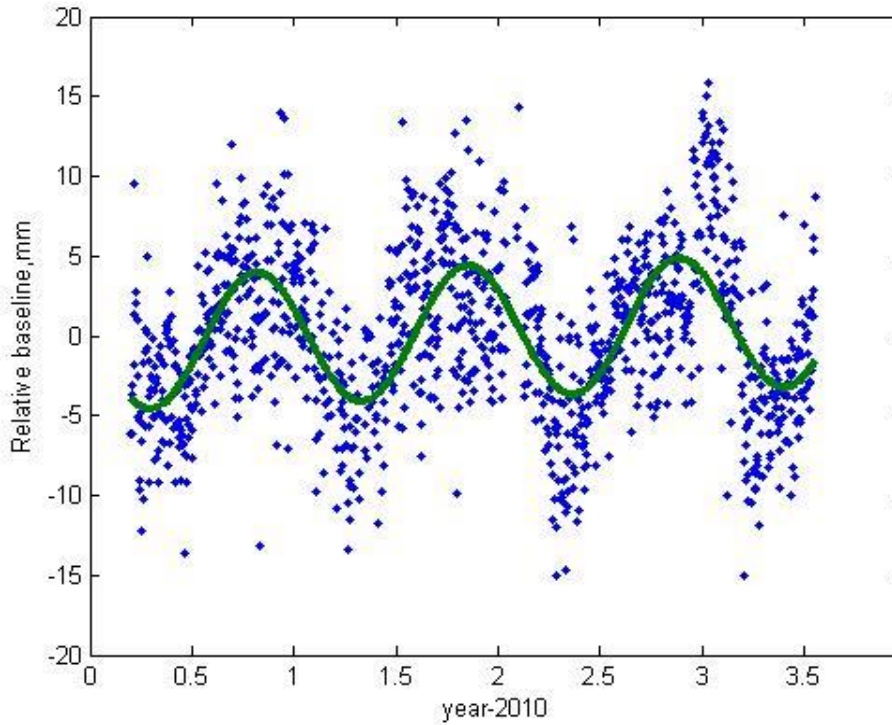


GNSS coordinate variations – WARK

WARK model residuals

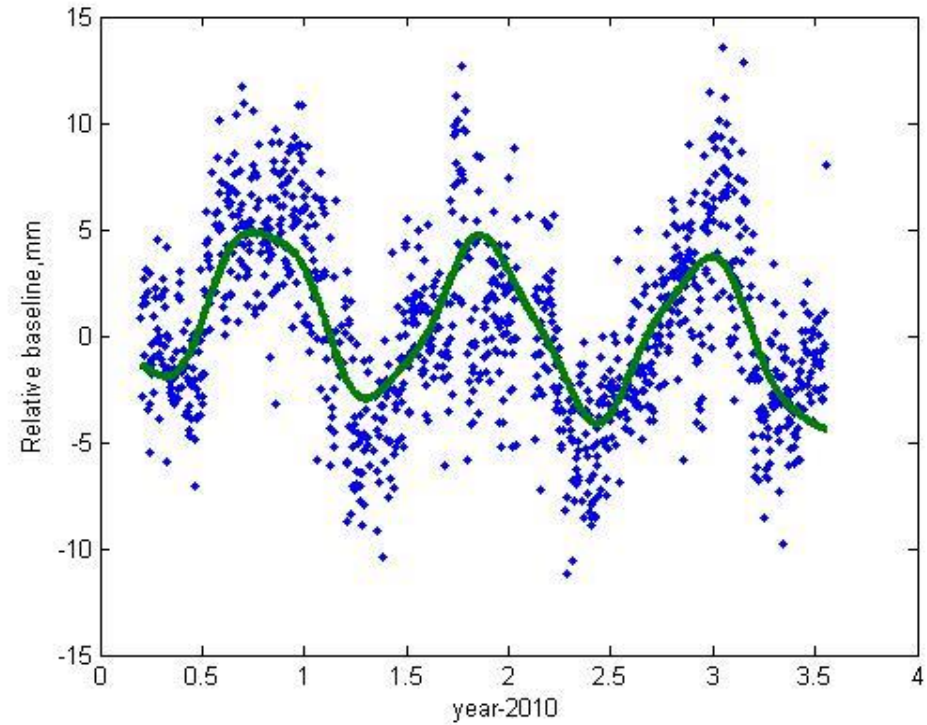
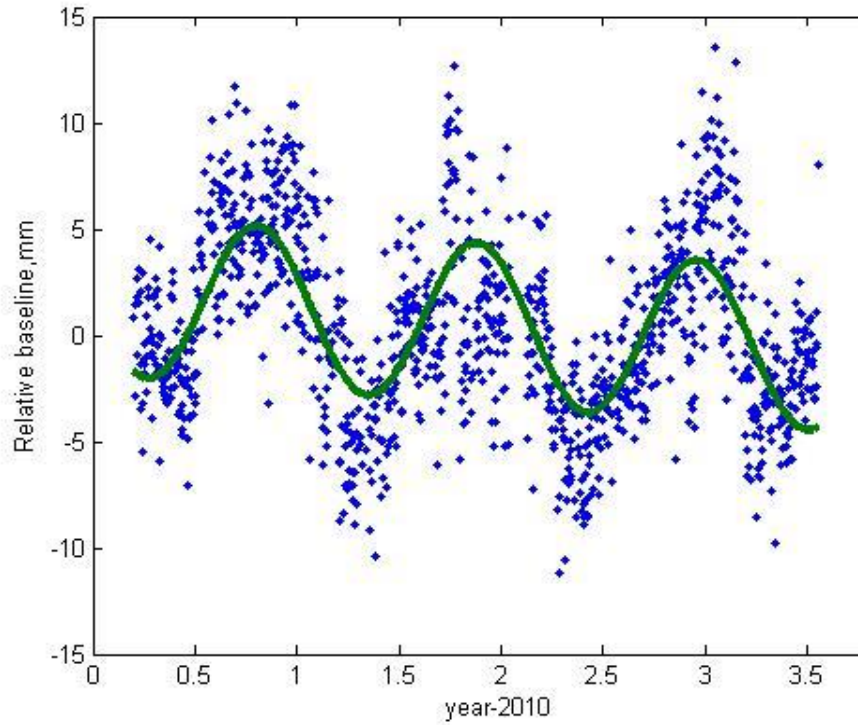


GNSS baseline variations



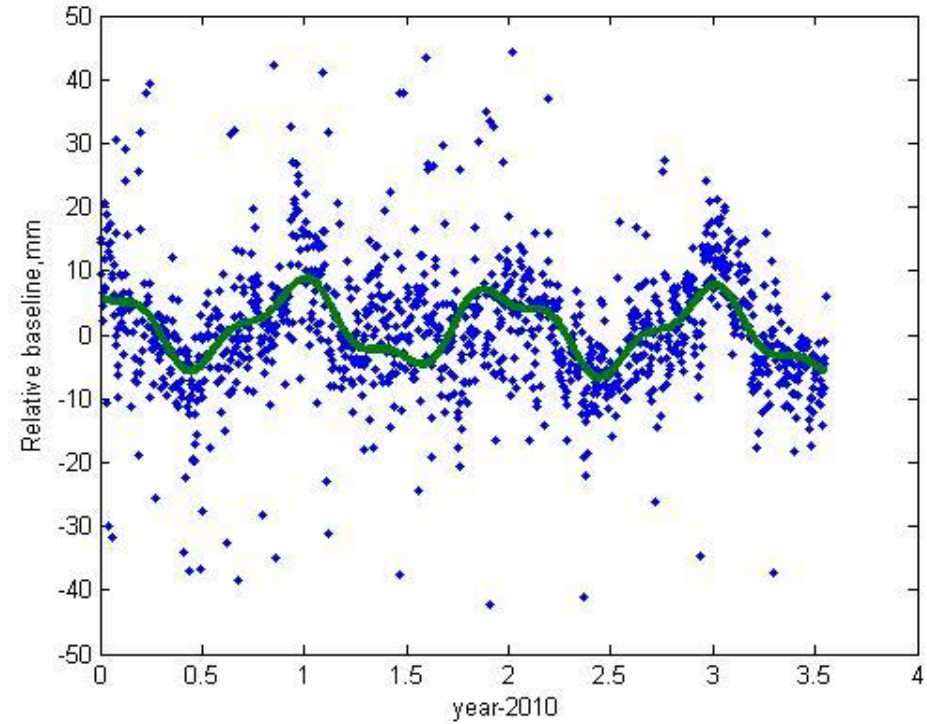
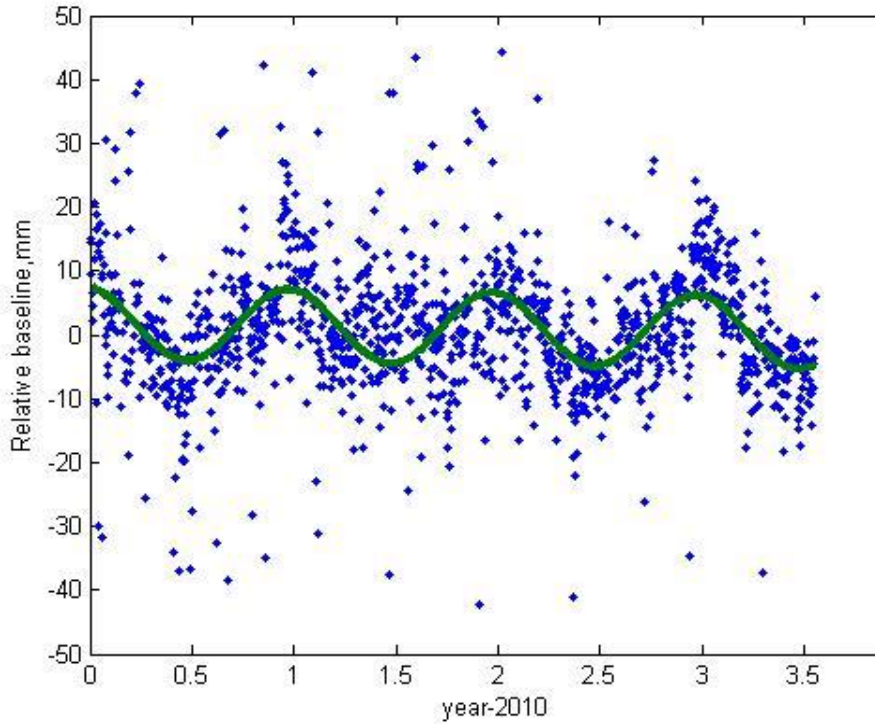
Ho-Ka GNSS baseline

GNSS baseline variations

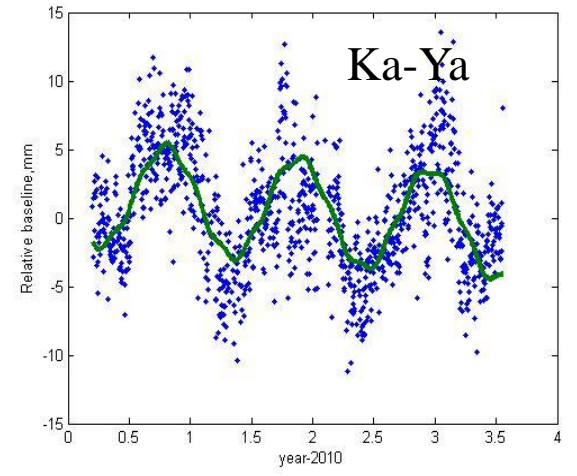
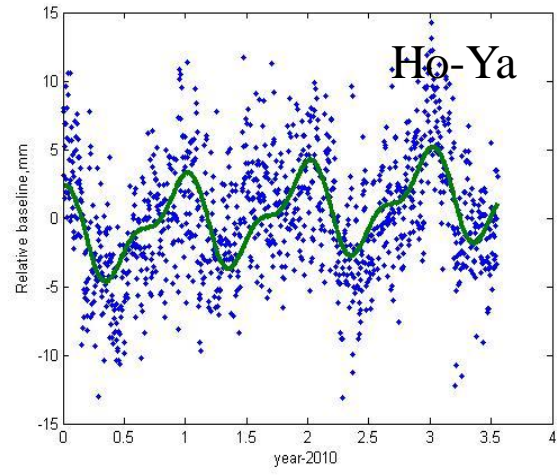
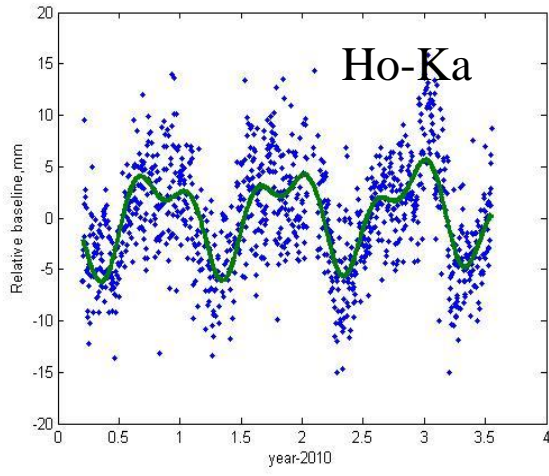
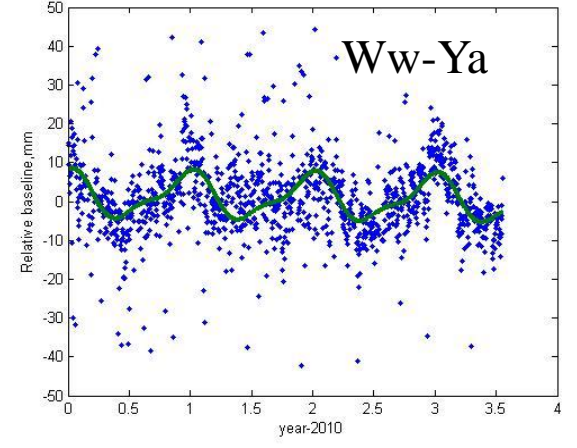
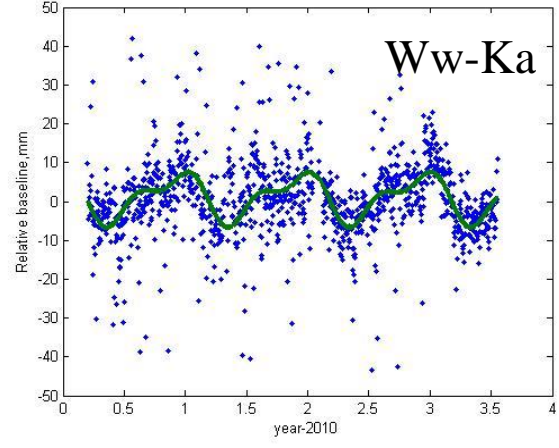
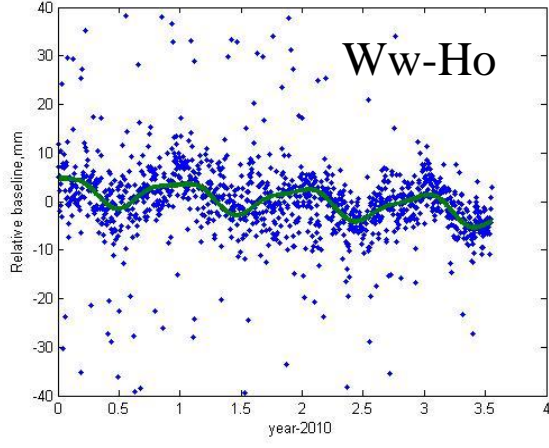


Ya-Ka GNSS baseline

GNSS baseline variations

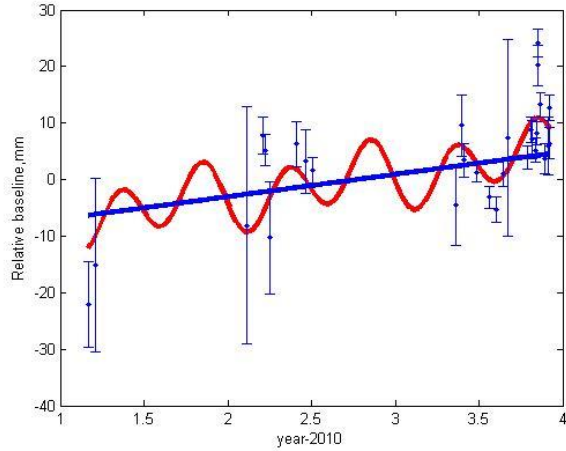


Ww-Ya GNSS baseline

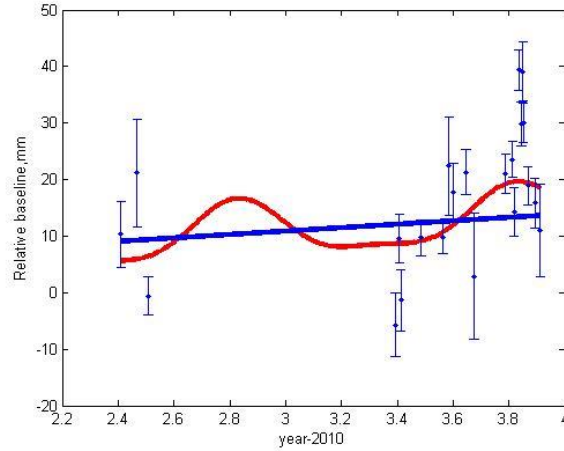


VLBI baseline variations

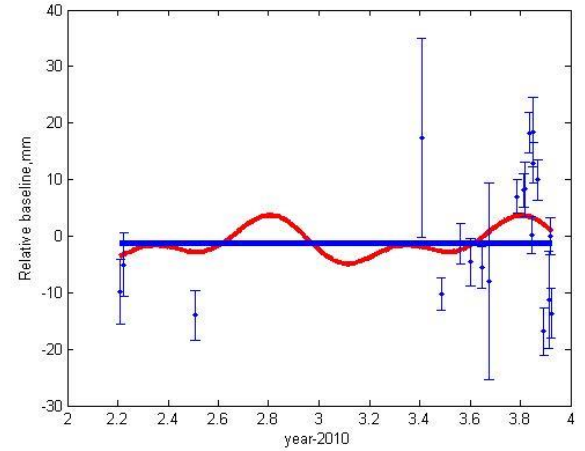
Ww-Ho



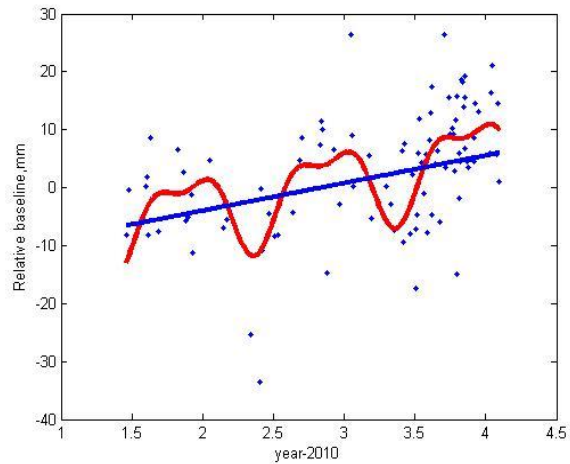
Ww-Ka



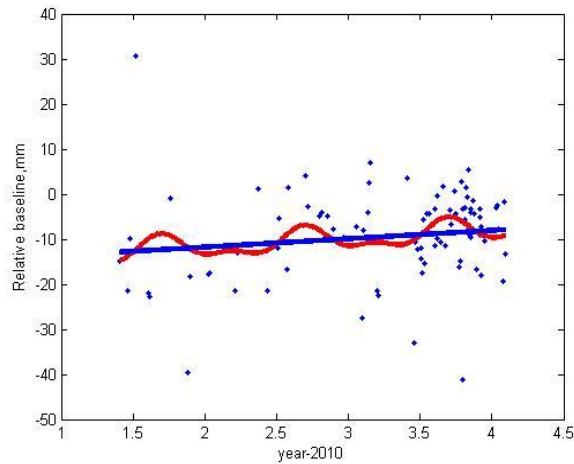
Ww-Ya



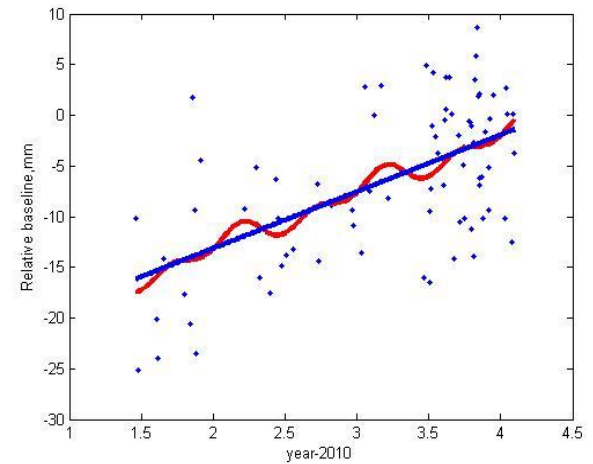
Ho-Ka



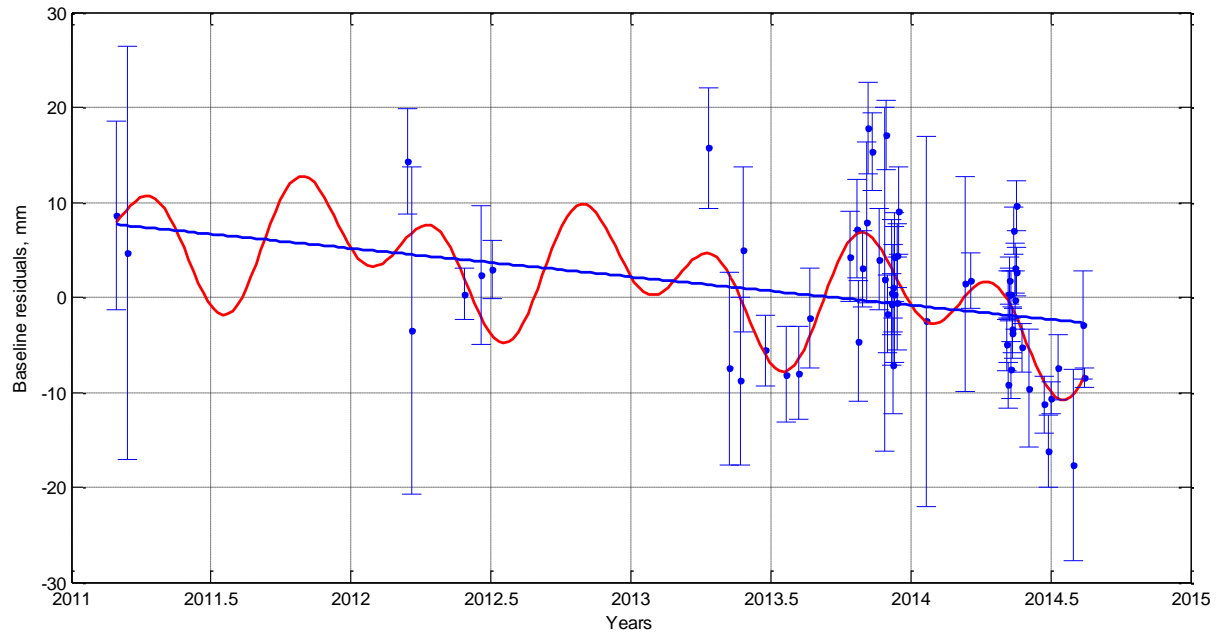
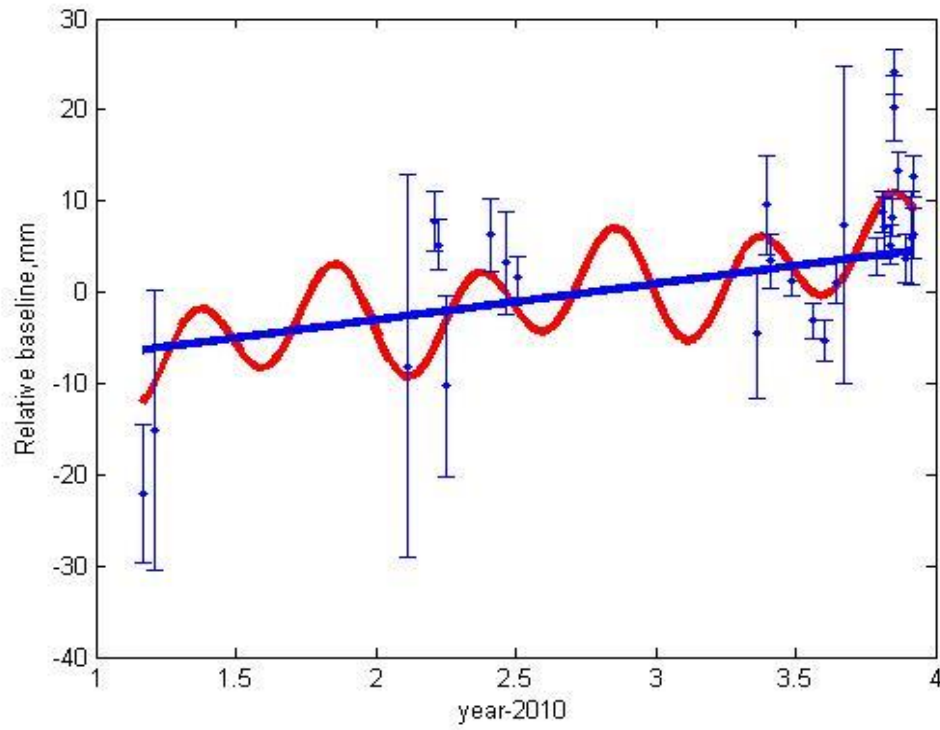
Ho-Ya



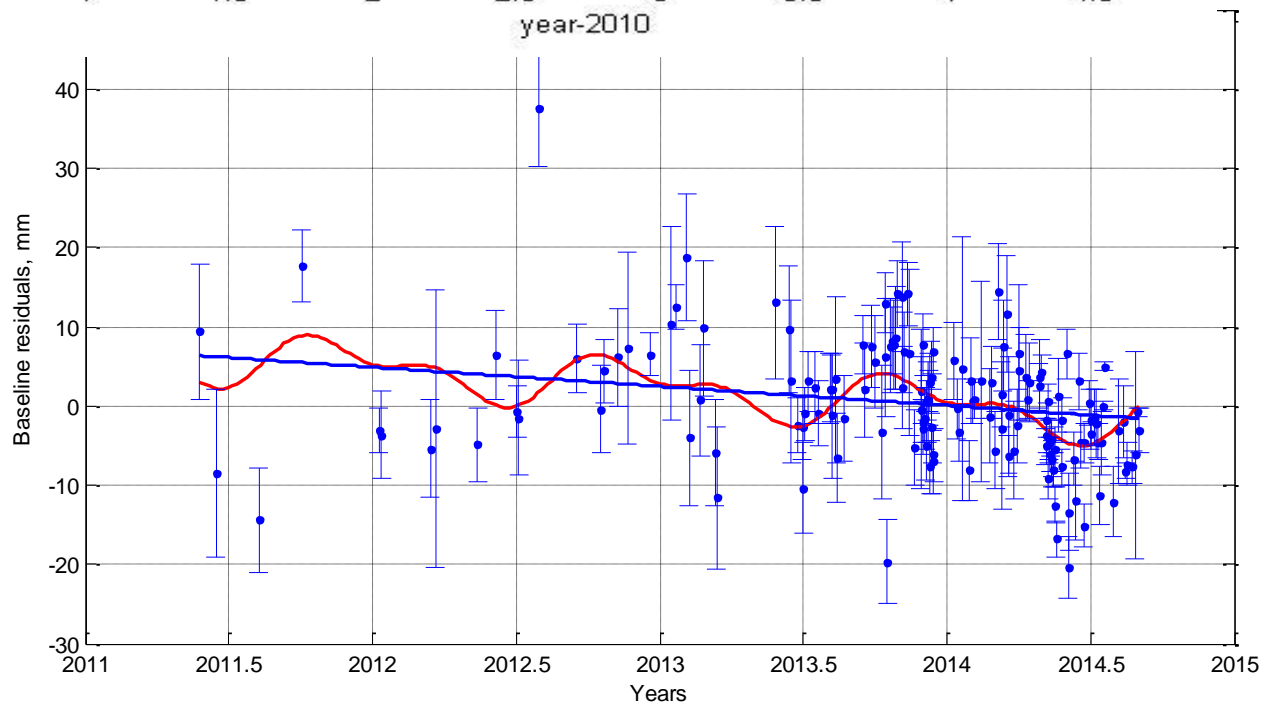
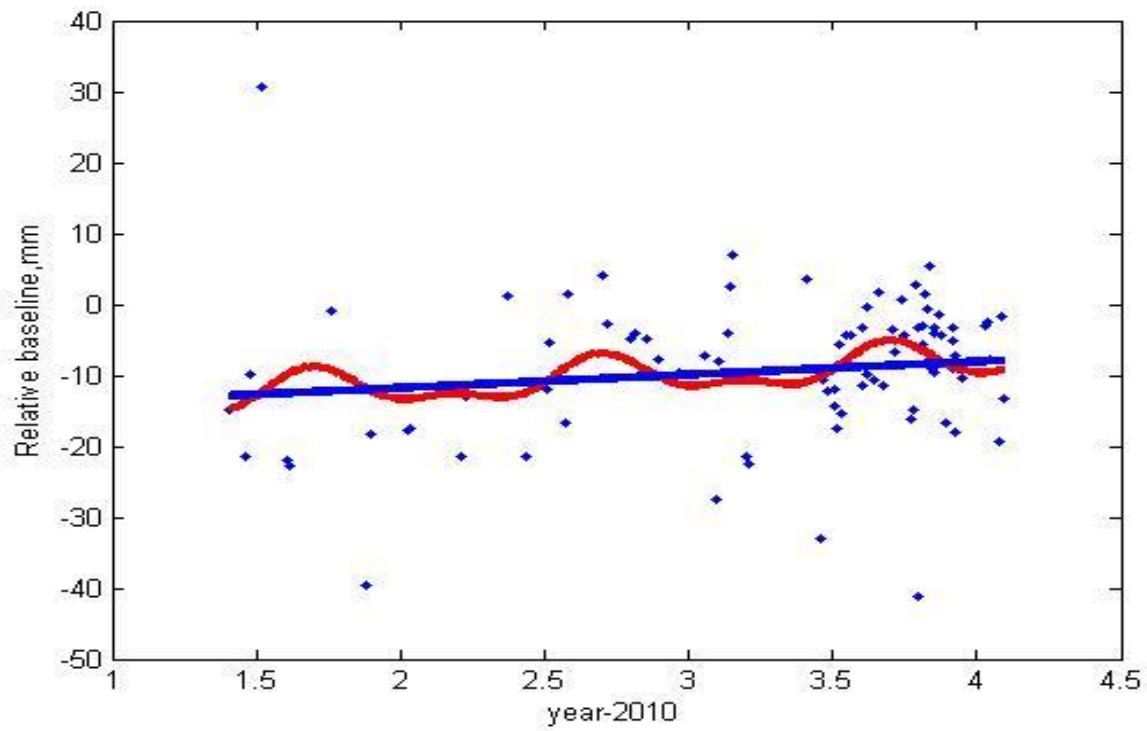
Ka-Ya



Ho-Ww



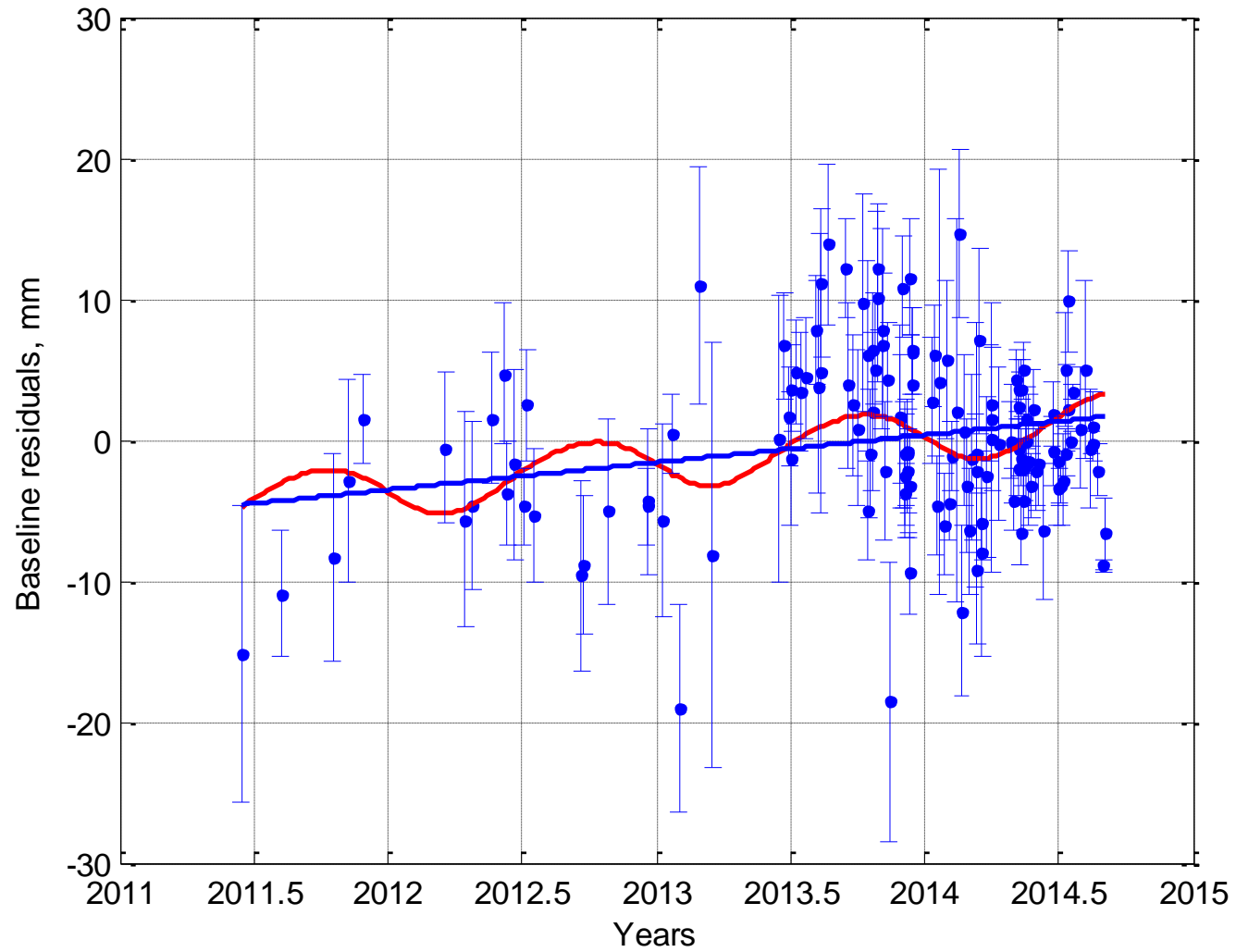
Ho-Ya



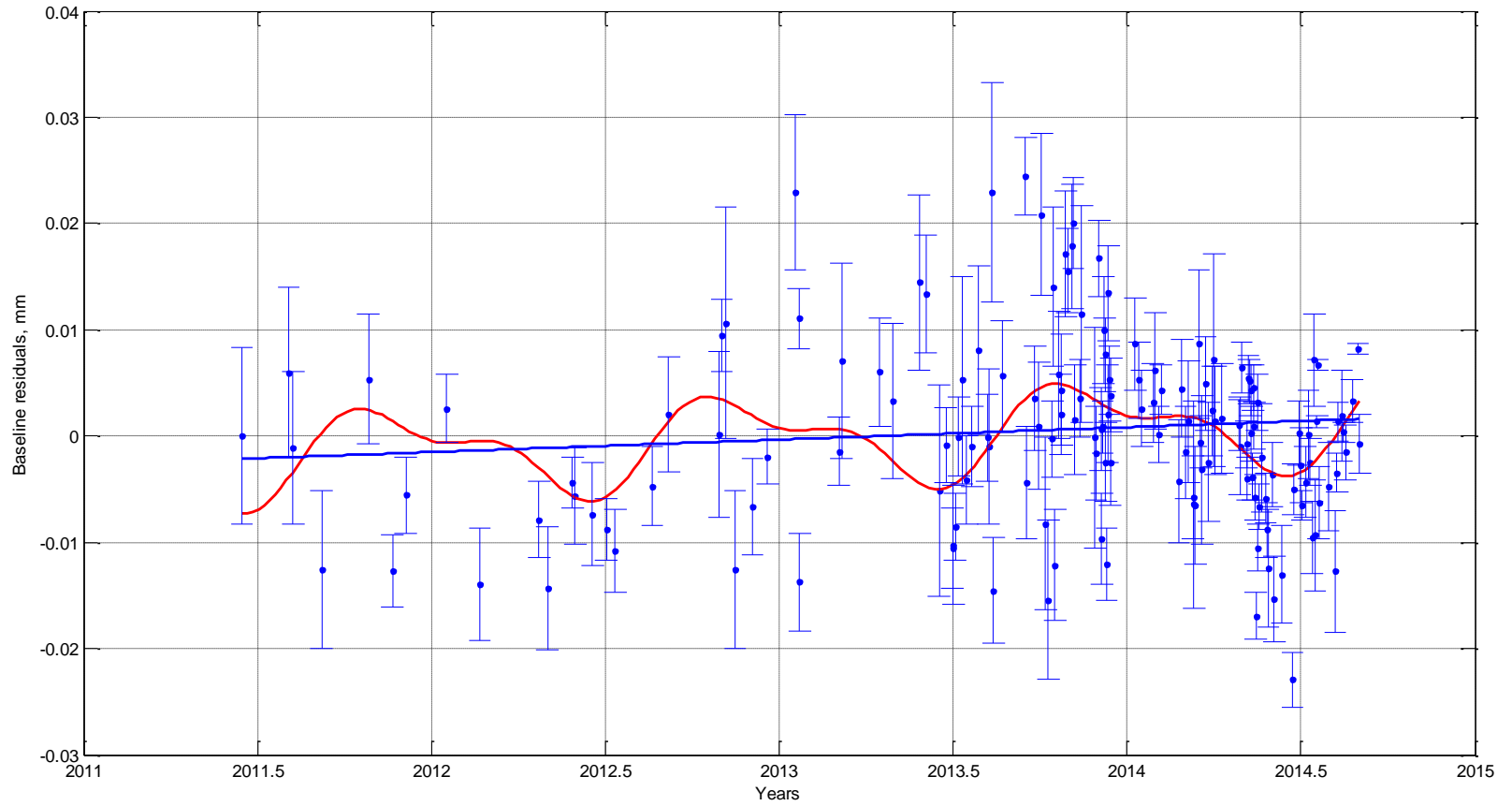
Coefficients	Table 1. GNSS baseline fitting parameters					
	WwHo	WwKa	WwYa	HoKa	HoYa	KaYa
x1(1)	2.75	5.47	5.49	4.03	2.86	3.50
x1(2)	-0.02	-0.10	-0.03	-0.13	-0.06	-0.12
x1(3)	2.78	1.17	1.73	-0.14	-1.36	2.25
x1(4)	-1.45	-0.02	-0.41	0.31	0.93	-0.78
x1(5)	0.59	2.82	1.94	2.18	1.37	1.20
x1(6)	-0.09	-0.93	-0.19	-0.17	0.06	-0.17
resnorm	7.5e+04	9.1e+04	1.1e+05	1.9e+04	1.7e+04	1.2e+04
stdev_resid	7.84	9.01	9.37	4.09	3.77	3.32
stdev_y	8.26	10.04	10.31	5.28	4.50	4.34

Coefficients	Table 2. VLBI baseline fitting parameters					
	WwHo	WwKa	WwYa	HoKa	HoYa	KaYa
x1(1)	2.13	4.79	2.87	6.65	2.34	0.82
x1(2)	-0.27	-0.17	-0.24	-0.13	-0.30	0.12
x1(3)	-10.96	1.78	-1.33	-13.49	-15.53	-24.35
x1(4)	3.95	3.03	0.02	4.76	1.86	5.60
x1(5)	5.13	-1.48	-2.20	-2.94	-1.23	-0.81
x1(6)	0.36	0.08	0.07	9.86	-0.06	-0.05
resnorm	1.1e+03	2.1e+03	2.1e+03	6.2e+03	8.5e+03	2.9e+03
stdev_resid,	6.36	9.64	10.21	8.15	9.94	6.22
stdev_y	9.59	12.06	11.33	10.45	10.23	7.72

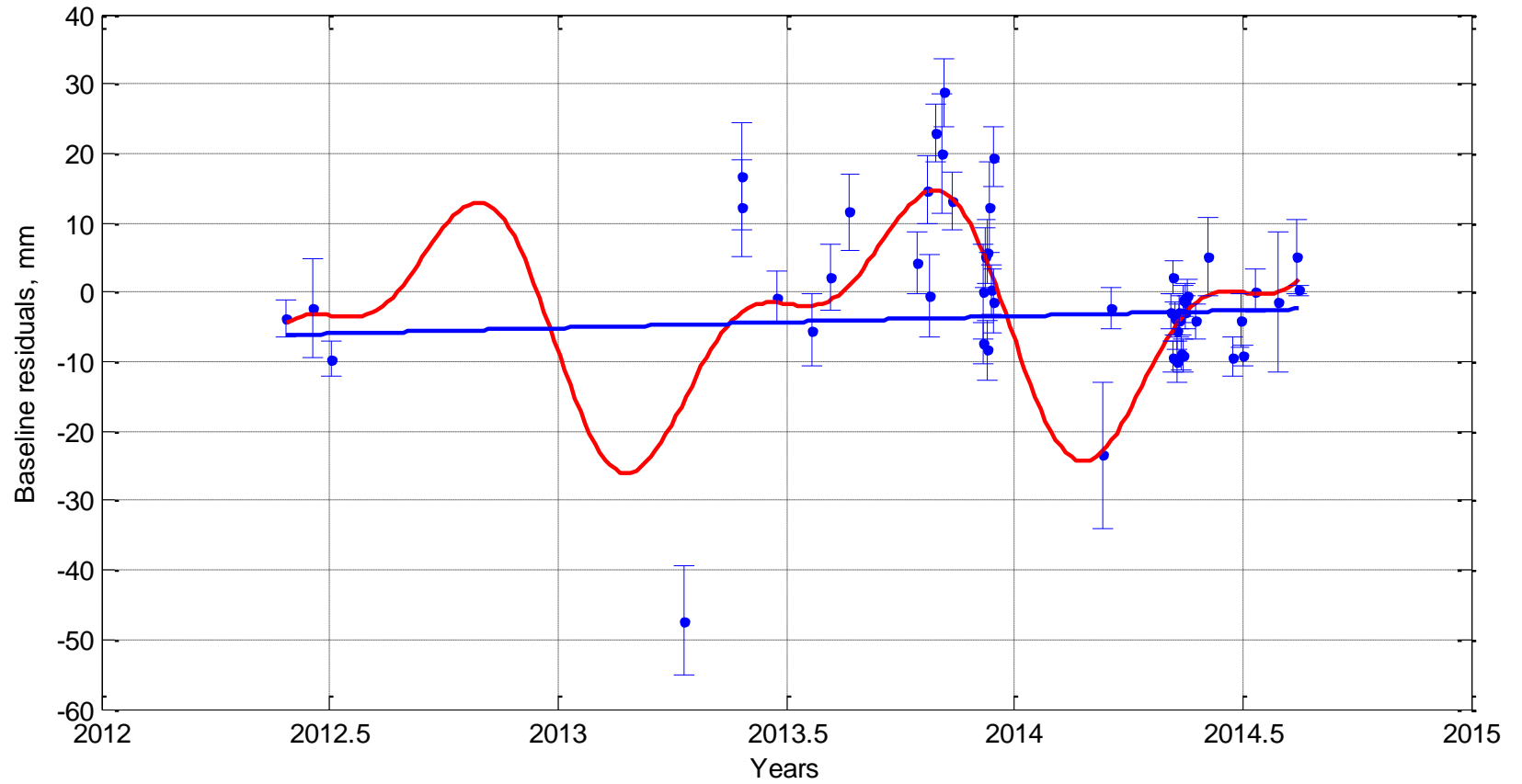
KaYa



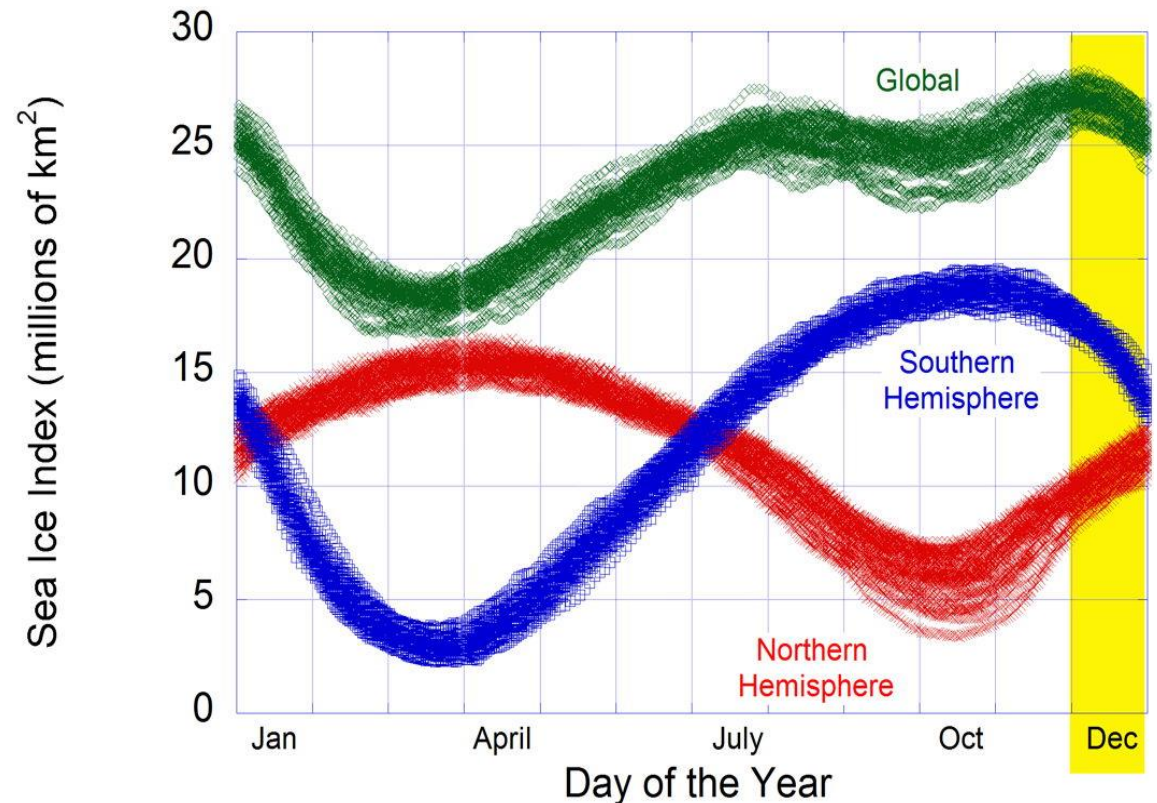
Ho-Ka



Ww-Ka



Sea ice extent



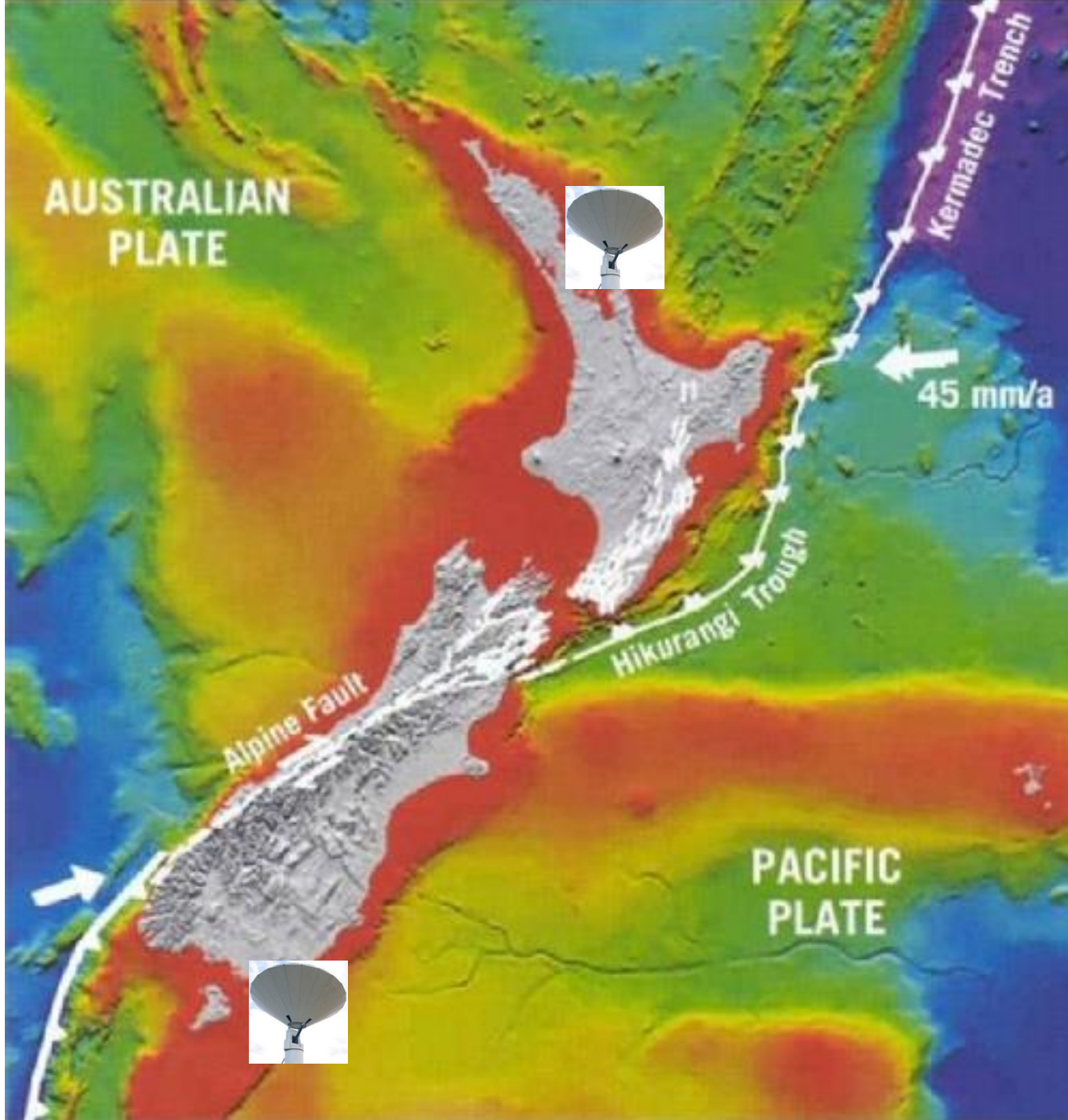
Data from US NSIDC showing the northern hemisphere, southern hemisphere and global sea ice extent plotted versus the day of the year. Based on 50 years observations



US Dept of State Geographer
© 2014 Google
Image U.S. Geological Survey
Data SIO, NOAA, U.S. Navy, NGA, GEBCO

Google earth



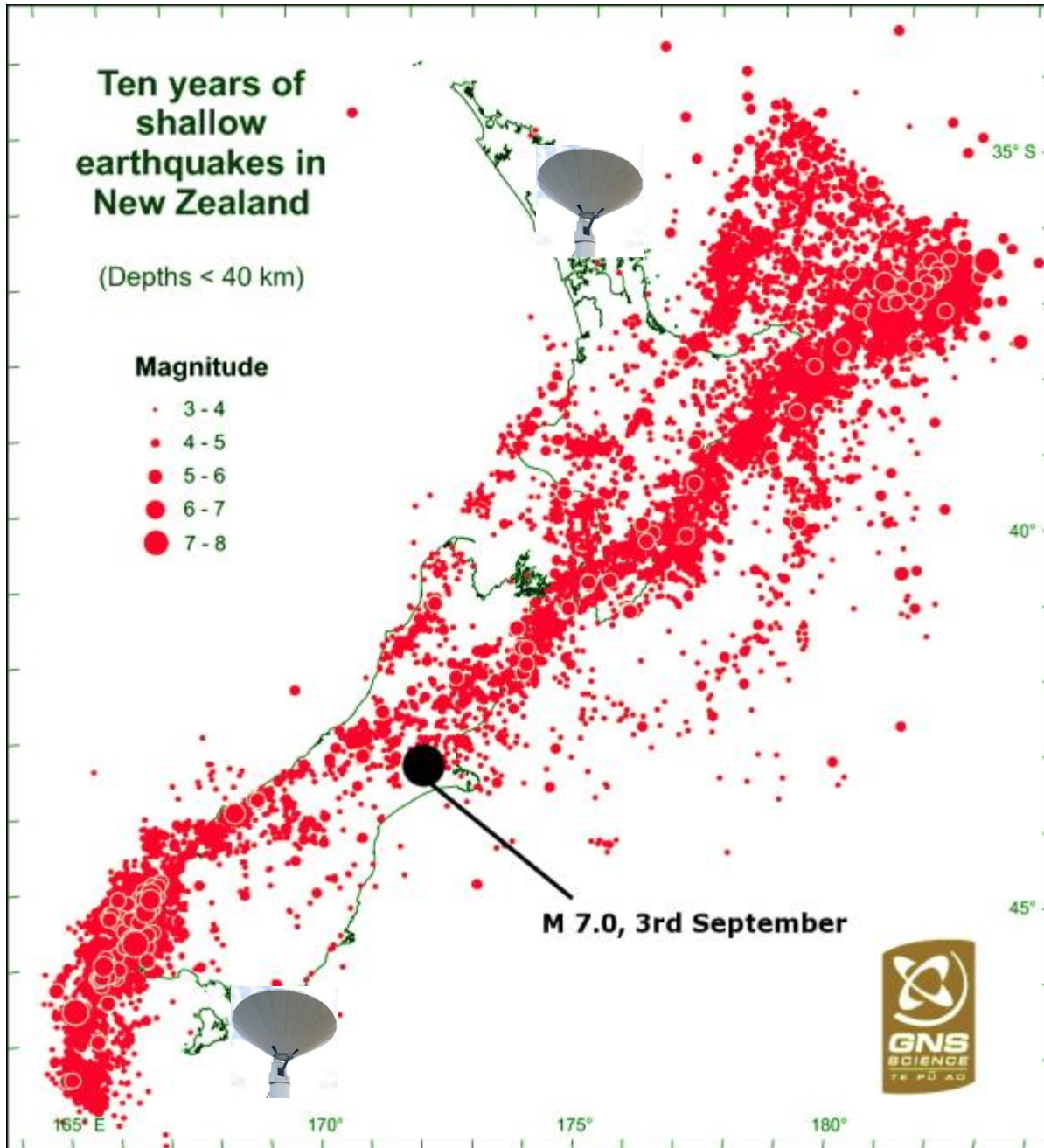


Ten years of shallow earthquakes in New Zealand

(Depths < 40 km)

Magnitude

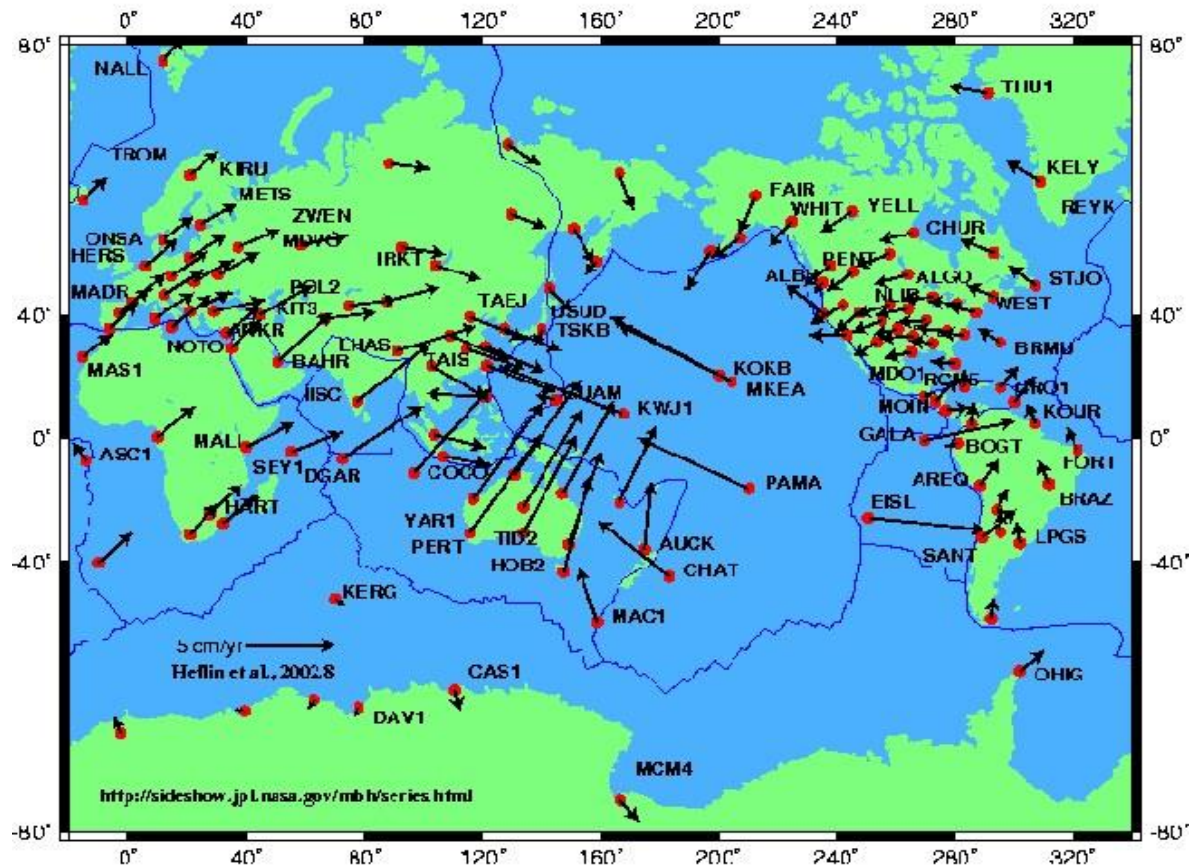
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- 4-5
- 5-6
- 6-7
- 7-8



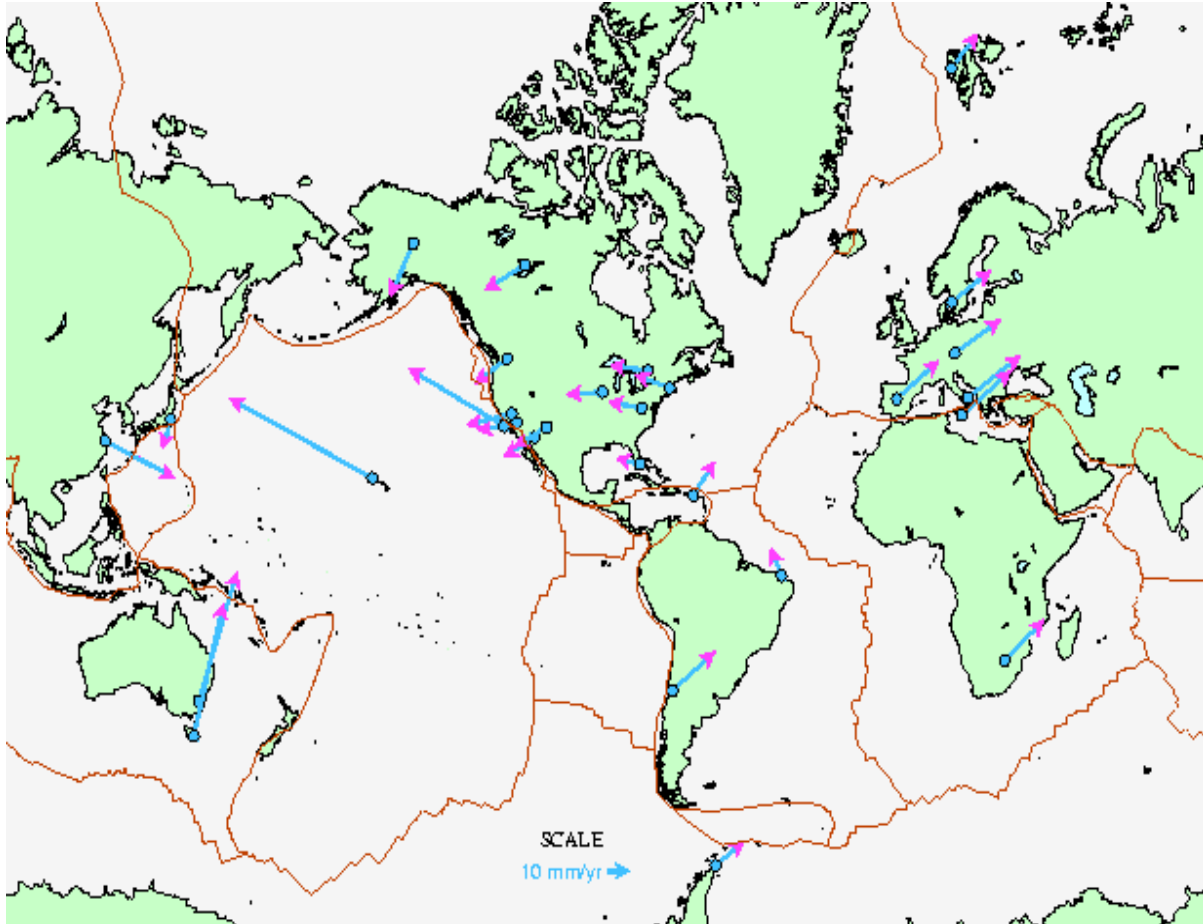
M 7.0, 3rd September



EOLSS - MANTLE DYNAMICS AND PLATE KINEMATICS

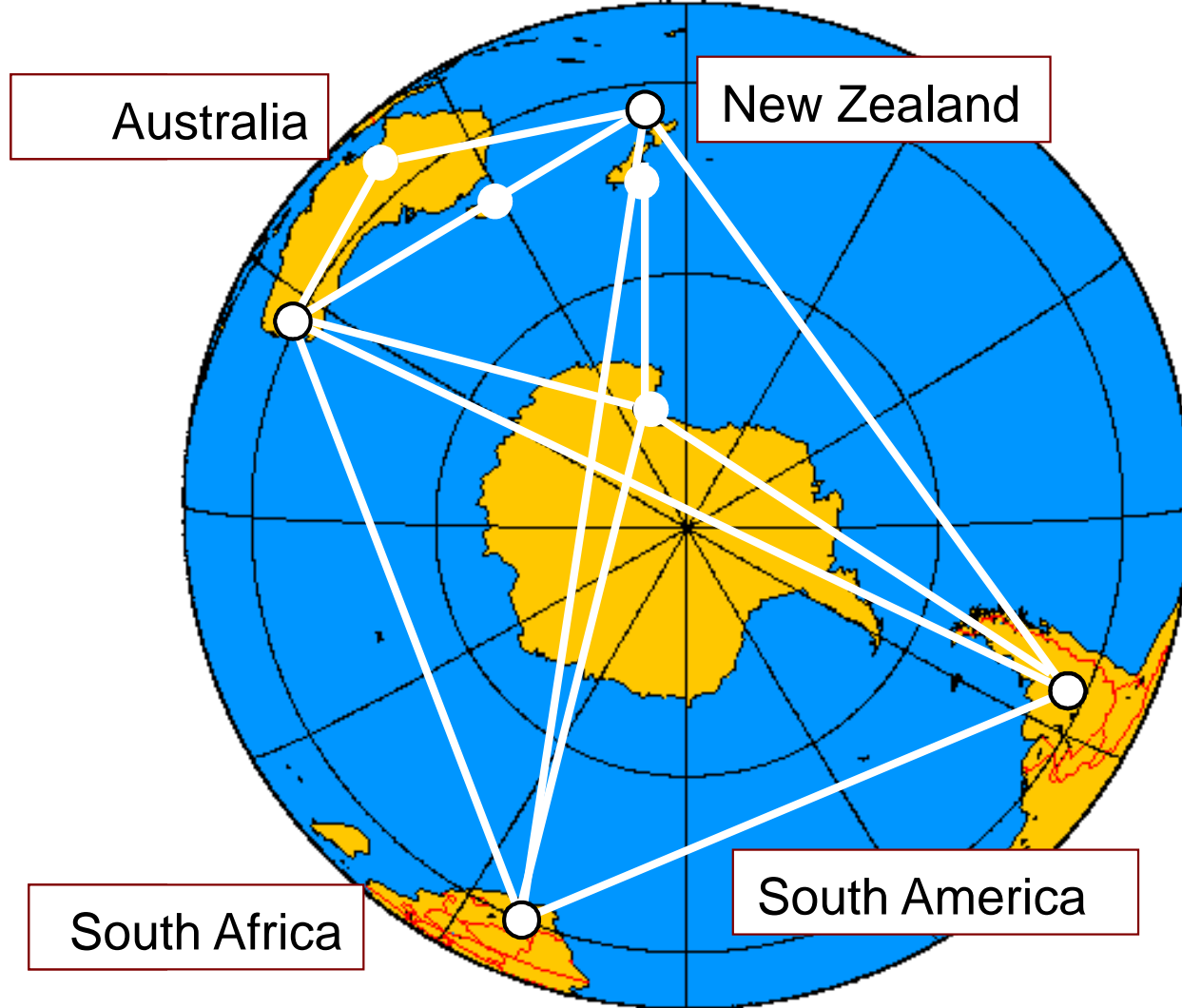


Motions of tectonic plates based on GNSS observations. The plot was compiled by Heflin et al., 2002 using the NASA database <http://sideshow.jpl.nasa.gov/post/series.html>



Annual movement of radio telescopes, measured by geodetic VLBI. The Hartebeesthoek telescope moves North-East at 2.5 cm per year. Tectonic plate boundaries are shown as brown lines.

<http://www.hartrao.ac.za/summary/sumeng.html>



Geographical distribution of GGOS stations – Australia, New Zealand, South Africa, South America. Locations of possible antennas in NZ South Island and in NZ Scott Base (Antarctica) are indicated. Projection: polar orthographic.

Thank you !